



祝
ノーベル賞受賞
山中伸弥氏
大阪教育大学附属天王寺中学校
中29
高23

祝
我らが先輩 山中伸弥先生
ノーベル賞受賞
おめでとうございます

祝
国際化学オリンピック銀賞受賞
高三遊谷亮太君

大阪教育大学附属天王寺中学校
大阪教育大学附属天王寺中学校

A Noble Prize Winner
No teacher supervise in exam classes
Crazy 30hours walking event etc.

Two Japanese novel prize winners

Shinya Yamanaka **山中伸弥**

The 2012 Nobel Prize for
Physiology or Medicine for
the discovery of iPS cells.

He is a graduate of Tennoji High school
attached to Osaka-Kyoiku University



Nobel Prizes and Laureates

Chemistry Prizes ▼ < 2002 >

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▶ [Kurt Wüthrich](#)

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The Nobel Prize in Chemistry 2002

John B. Fenn, Koichi Tanaka, Kurt Wüthrich

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The Nobel Prize in Chemistry 2002



John B. Fenn
Prize share: 1/4



Koichi Tanaka
Prize share: 1/4



Kurt Wüthrich
Prize share: 1/2

The Nobel Prize in Chemistry 2002 was awarded *"for the development of methods for identification and structure analyses of biological macromolecules"* with one half jointly to John B. Fenn and Koichi Tanaka *"for their development of soft desorption ionisation methods for mass spectrometric analyses of biological macromolecules"* and the other half to Kurt Wüthrich *"for his development of nuclear magnetic resonance spectroscopy for determining the three-dimensional structure of biological macromolecules in solution"*.

Photos: Copyright © The Nobel Foundation

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Nobel Women

Koichi Tanaka (田中 耕一 Tanaka Kōichi, born August 3, 1959) is a Japanese engineer who shared the Nobel Prize in Chemistry in 2002 for developing a novel method for **mass spectrometric analyses** of biological macromolecules with John Bennett Fenn and Kurt Wüthrich (the latter for work in NMR spectroscopy)



THE PRIZE STORY

The Prize Story

The Road to Success



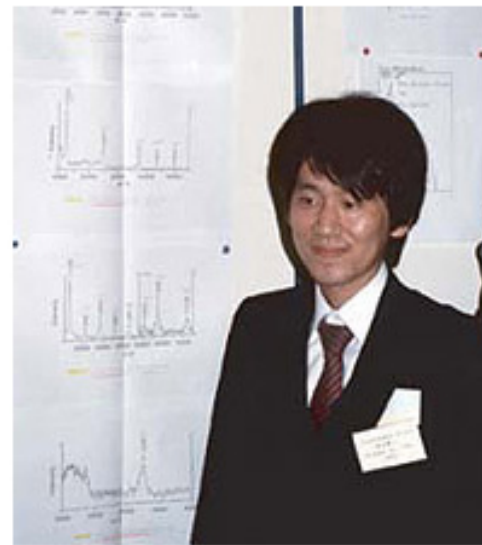
Tanaka answering questions from reporters at the press conference at 9:00pm, October 9th, 2002.

'83: A Quintet Team Is Created

Back in 1980 Shimadzu Corporation newly established a Central Research Laboratory comprising three research groups (C for chemistry, M for mechanics and E for electronics), each staffed by over ten researchers. Sometime later Koichi Tanaka was to find himself affiliated to E-group where one of the research themes was "lasers" a leading technology at the time.

The Prize Story

The Road to Success



Koichi Tanaka in front of a poster-arranged announcement at the 2nd Japan-China Joint Symposium on Mass Spectrometry held in Takarazuka, Japan in 1987 (Photo by courtesy of Dr. Cotter)

Meeting The "World" In '87

Koichi Tanaka first met the "world" in September 1987 at the 2nd Japan-China Joint Symposium on Mass Spectrometry held in Takarazuka, Japan. It was to this event that Dr. Robert Cotter a leading authority, even at that time, in the mass spectrometry field had been invited. He stated in his address that it was more than likely impossible to detect macromolecules using laser ionization mass spectrometry. After the address, Tanaka, who was present at the symposium on behalf of the research team to organize a poster-arranged presentation, approached Dr. Cotter and informed him of their research results showing the amazing data in which detection reached a mass number of 72,000. This data had a profound impact on Dr. Cotter, who promptly took report copies back to the USA and, in joint co-operation with Dr. Catherine Fenselau, a biological macromolecule researcher, introduced Shimadzu research results to many researchers in Europe and USA.

DETECTION OF HIGH MASS MOLECULES BY LASER DESORPTION TIME-OF-FLIGHT MASS SPECTROMETRY

Koichi Tanaka, Yutaka Ido, Satoshi Akita, Yoshikazu Yoshida and Tamio Yoshida

Central Research Laboratory, Shimadzu Corporation, 1 Nisinokyo-Kuwabaracho, Nakagyo-ku, Kyoto 604, Japan

[Introduction]

The laser desorption time-of-flight mass spectrometer has been developed in order to analyze non-volatile, thermally labile and high mass organic molecules. In this spectrometer we have made improvements on all stages of mass spectrometer (ion source, mass separation, detector, electronics).

[Equipment]

The construction of the laser desorption time-of-flight mass analyzer is shown in Fig.1. Figure 2 shows the block diagram of TOF spectrum measurement system.

-- Ion source --

N_2 laser (Wavelength:337nm, Pulse width:about 15nsec, Pulse energy:4mJ Max.) was used for ionization. "Rapid heating" [1] is achieved by irradiating pulsed laser on sample surface. As for sample preparation, "Ultra fine metal powder (UFP) and glycerol matrix method" was found to be very effective for increasing the yield of high mass molecular ions, and decreasing the yield of fragment ions[2].

In comparison with Bulk, UFP has the following features

- High photo-absorption
- Low heat capacity
- Extremely large surface area per unit volume

This UFP matrix method seemed to enhance the speed of heating even further.

-- Mass separation --

Generally, TOF-MS has the following characteristics

- Very high transmission
- Measurement time of less than a few hundreds μ sec

- Unlimited mass range
- Low mass resolution

A new gradient-electric field ion reflector for a time-of-flight mass spectrometer has been developed in order to improve mass spectral resolution by energy focusing [3]. In the TOF mass spectrometer consisting of a free ion drift region and a new ion reflector, the motion of the same m/z ions is quasi-single oscillation of the same period. Therefore, the flight-times of the same m/z ions are focused to a constant even if the initial kinetic energy of the emitted ions are scattered.

The TOF mass separation system was designed to permit easy switching between "Reflector type" ($V_e > V_0$) and "Linear type" ($V_e = 0$).

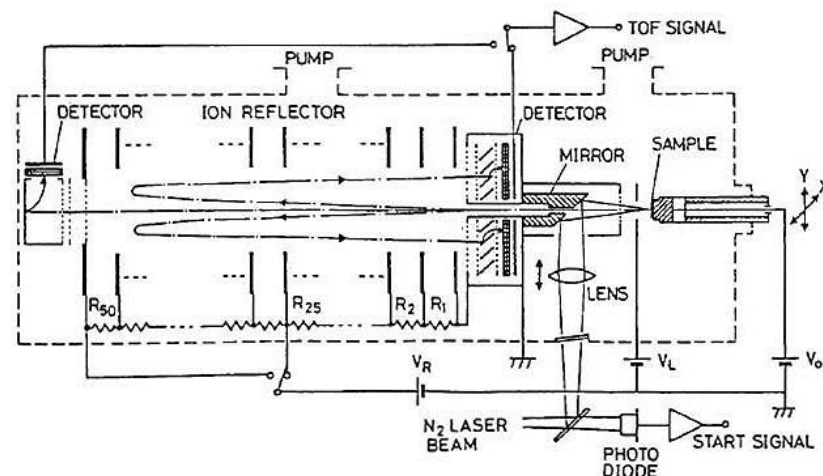


Fig. 1 Construction of the Laser Desorption TOF Mass Analyzer.

-- Detector --

Micro channel plate (MCP) or secondary electron multiplier (SEM) is usually used to detect ions, electrons or photons. Ions of larger m/z generally have low velocities in TOF-MS. So the detection sensitivity of MCP has a tendency to decrease in higher mass regions.

Higher detection sensitivity for high mass ions was



Early Earth and **South Africa Geology**

Yoshio Okamoto

Geoscience-English lecture
29th Nov. 2016



Why South Africa?

As a Geological wonderland

**Oldest rocks (3.5Ga<): Canada, Greenland, Western Australia:
most accessible locality!**

⇒ In Japan, no rock of this era.

Economic Ores: Gold, Diamond, Platinum etc.

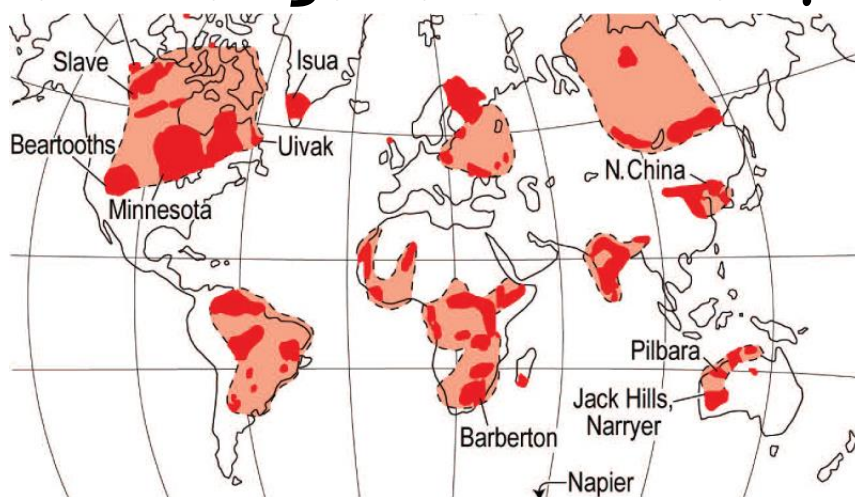
Gondwana homeland of continental drift

Oldest magma intrusion: "Bushveld igneous intrusion"

Meteor impacts Crater

Banded Iron Formation (BIF)

Ancient ice age remnants etc.



**Global distribution of Archean rocks
in modern continents.**

**Known (red), suspected (pink). Areas
with rocks or zircons older than 3.6
billion years are labelled by name.**

http://www.earthsciences.hku.hk/shmuseum/earth_evo_03_archean_intro.php

My Visits (Three-times)

IAGOD Geological Field Trip (Johannesburg to Pretoria): Aug. 2002: Big Five economic mines: Gold, Platinum, Chromium, Diamond, Iron, Coal

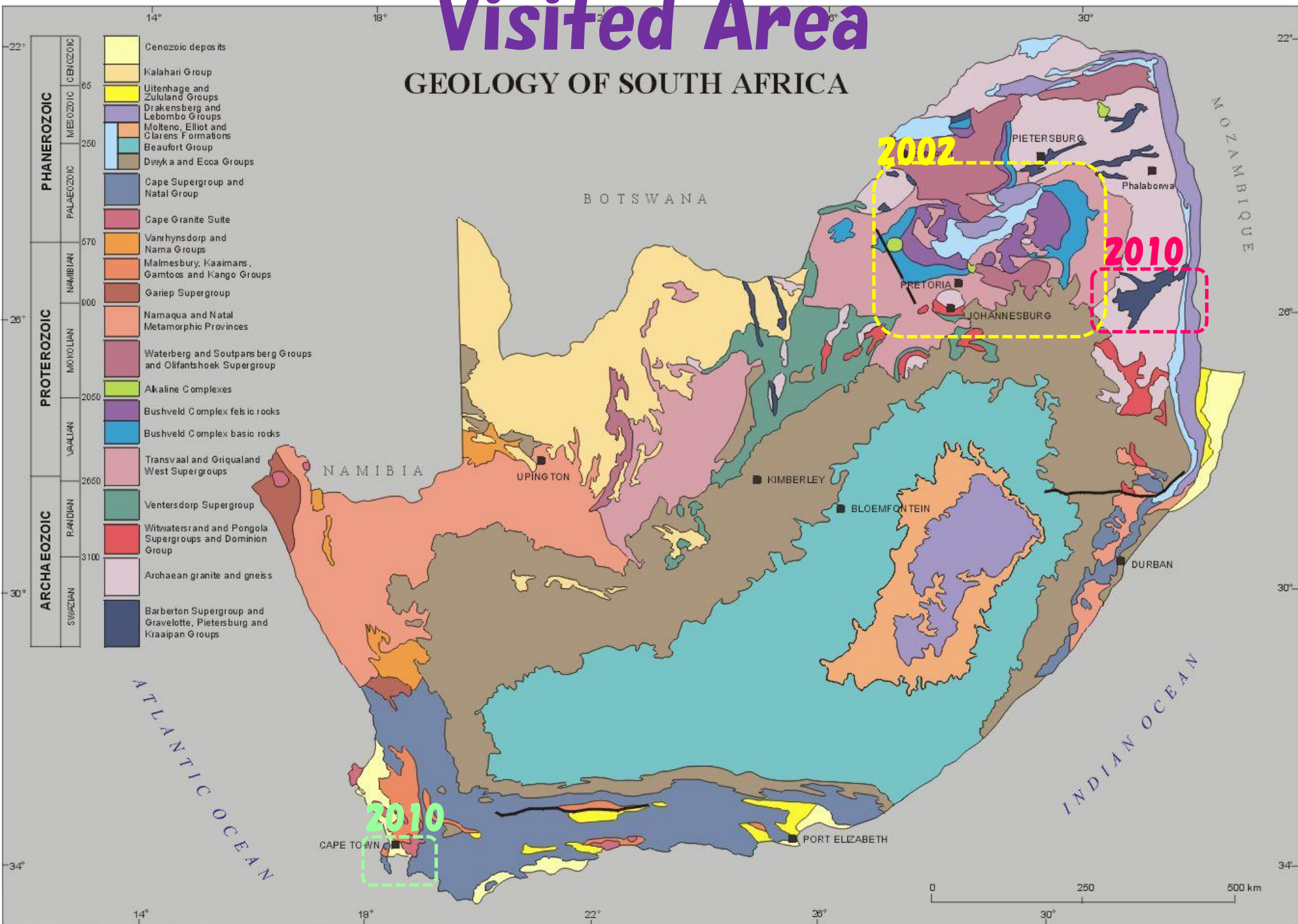
GeoSciEdVI Conference(Johannesburg) and field trip: Aug.-Sep. 2010: Barberton and Kruger National Park, Tswaing Meteor Impact Crater

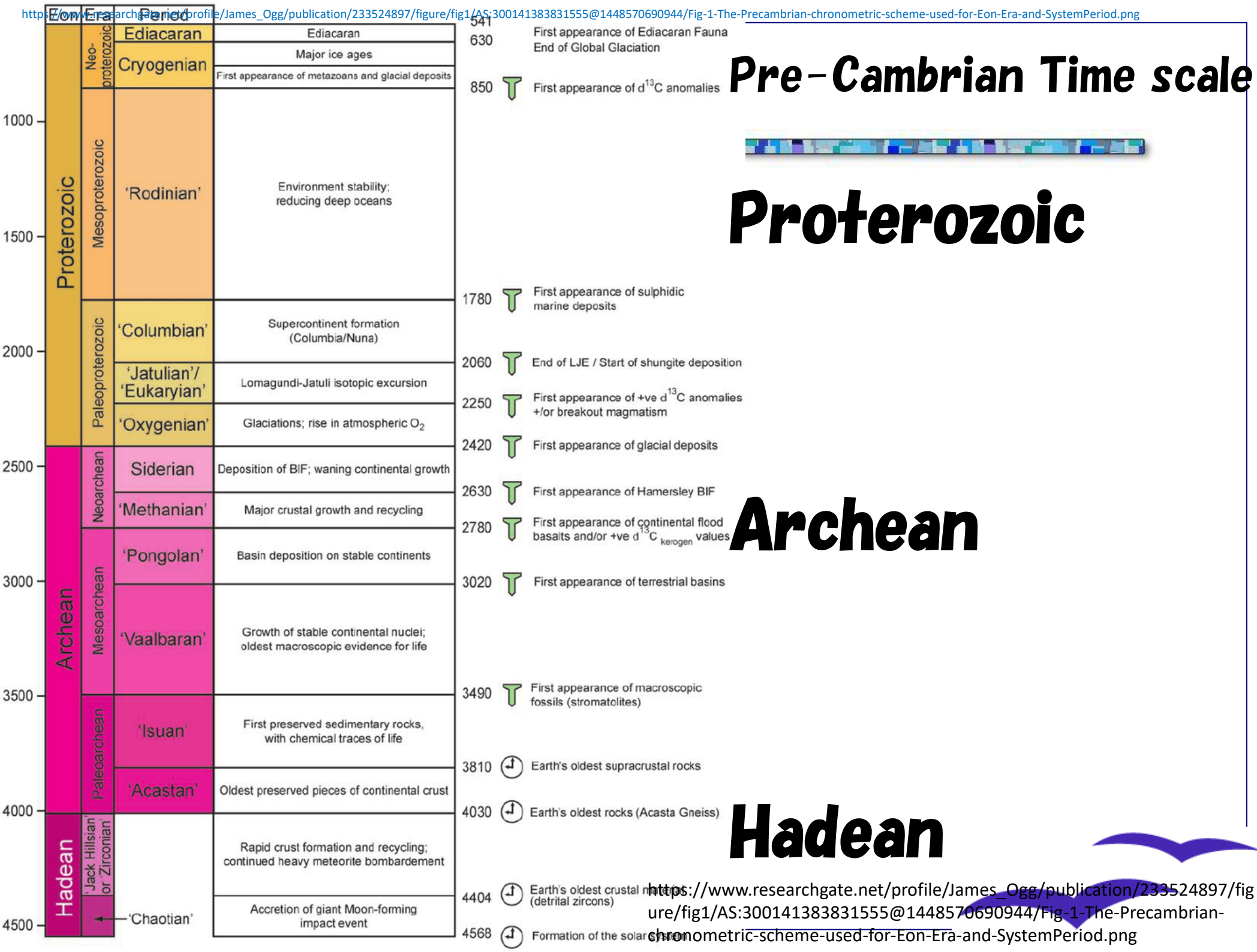
IGC 35 Conference and Field trip: Aug-Sep. 2016: Robben Island, Cape of Good Hope, Table Mountain



Visited Area

GEOLOGY OF SOUTH AFRICA





Pre-Cambrian Time scale

Proterozoic

Archean

Hadean

Absolute years

**Gy = Ga = Billion years ago =
1000000000 years**

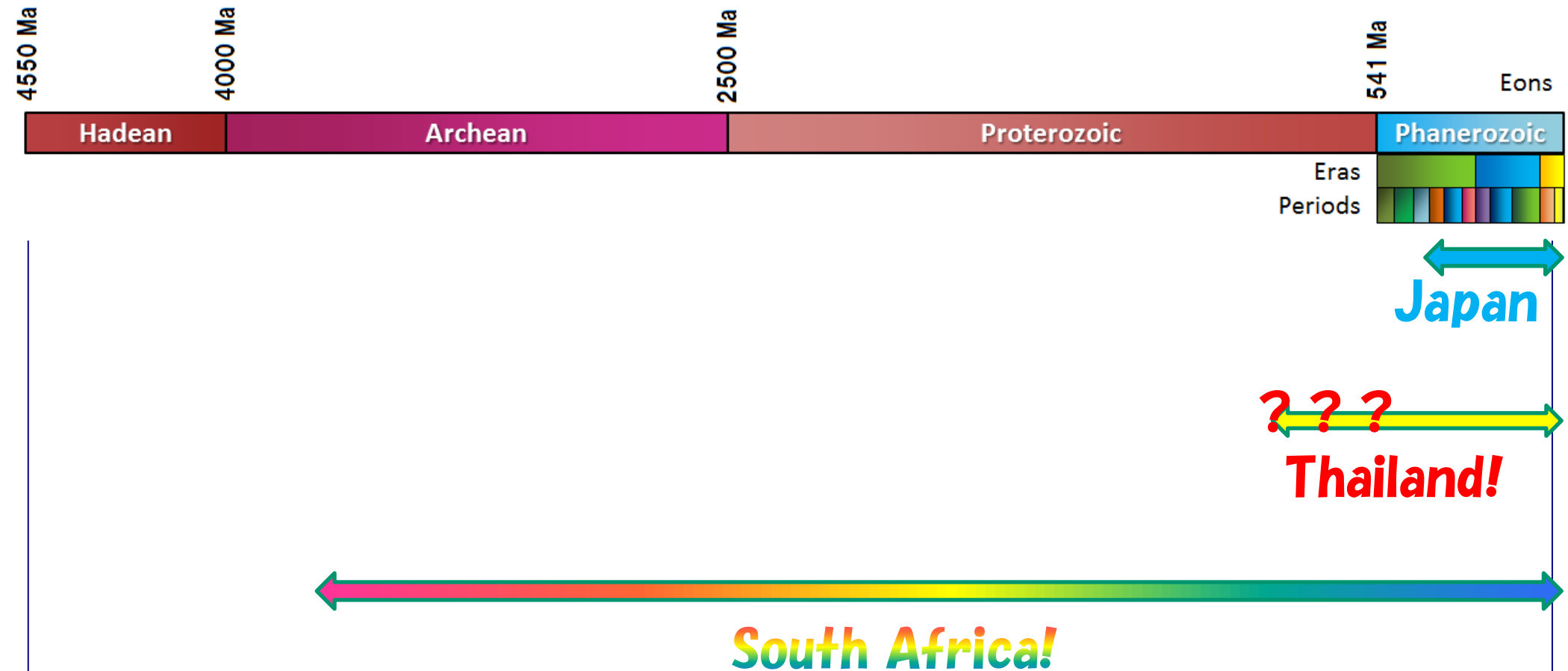
**My = Ma = Million years ago =
1000000 years**

3.0Ga = 3000Ma = 3000000000 years

**Birth of Earth = 4.6 Ga = 4600 Ma =
4600000000 years**



Thailand, Japan vs. South Africa on Geological Time scale



Why recently is the early earth so revealed.

After 1990' s: A radiometric dating tool is developed:

“SHRIMP, Sensitive high-resolution ion microprobe”

-> 20 μm Zircon Pb/U, Pb/Pb

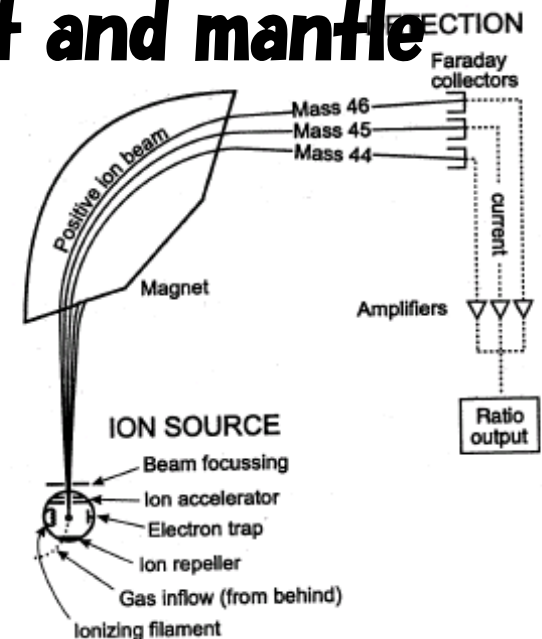
Also isotope ratio geochemical techniques are advanced:

$^{146}\text{Sm}-^{142}\text{Nd}$: $^{182}\text{Hf}-^{182}\text{W}$: $^{142}\text{Nd}/^{144}\text{Nd}$, $^{182}\text{W}/^{184}\text{W}$ ->

use for evolution of early earth crust and mantle system

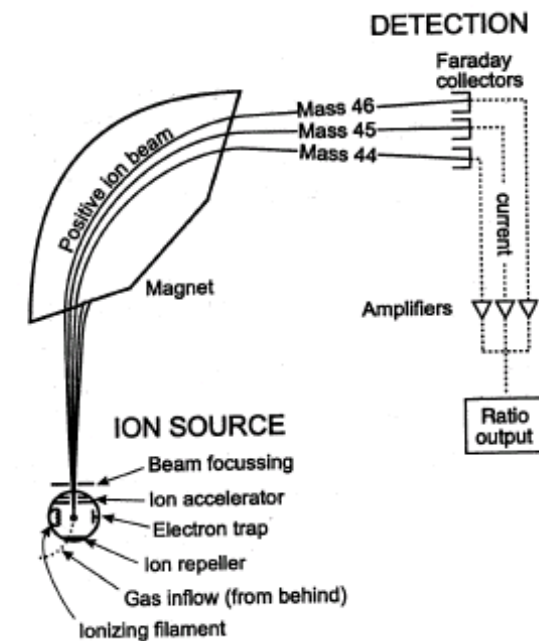
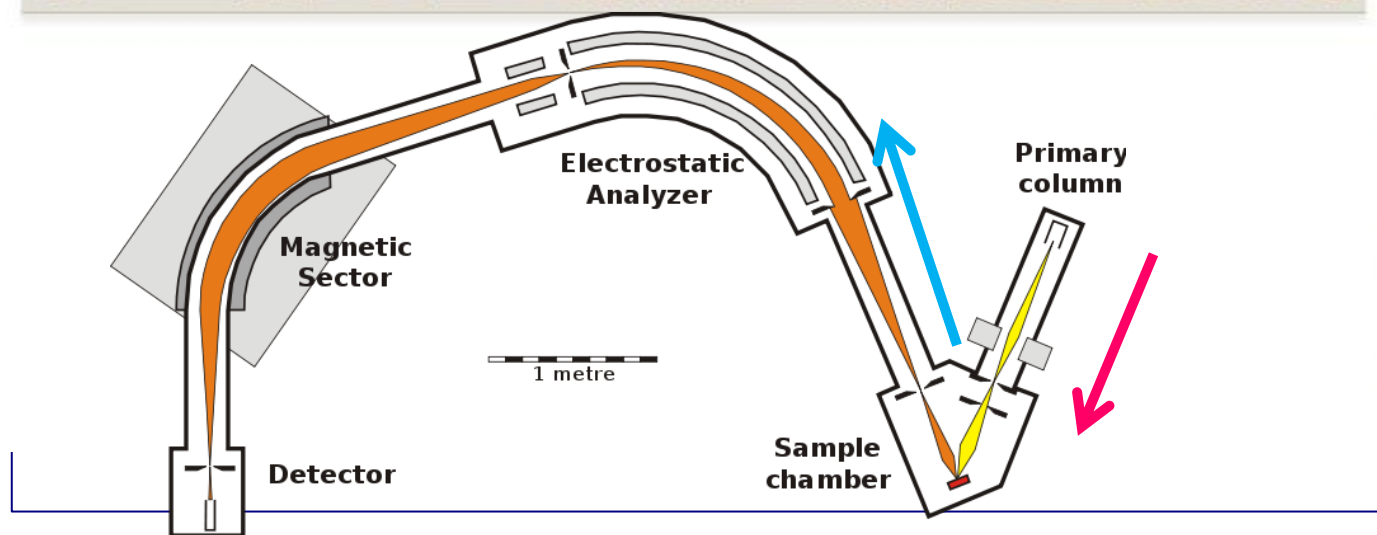
<Applying Mass Spectrometry>

A new window is opened for the early earth!





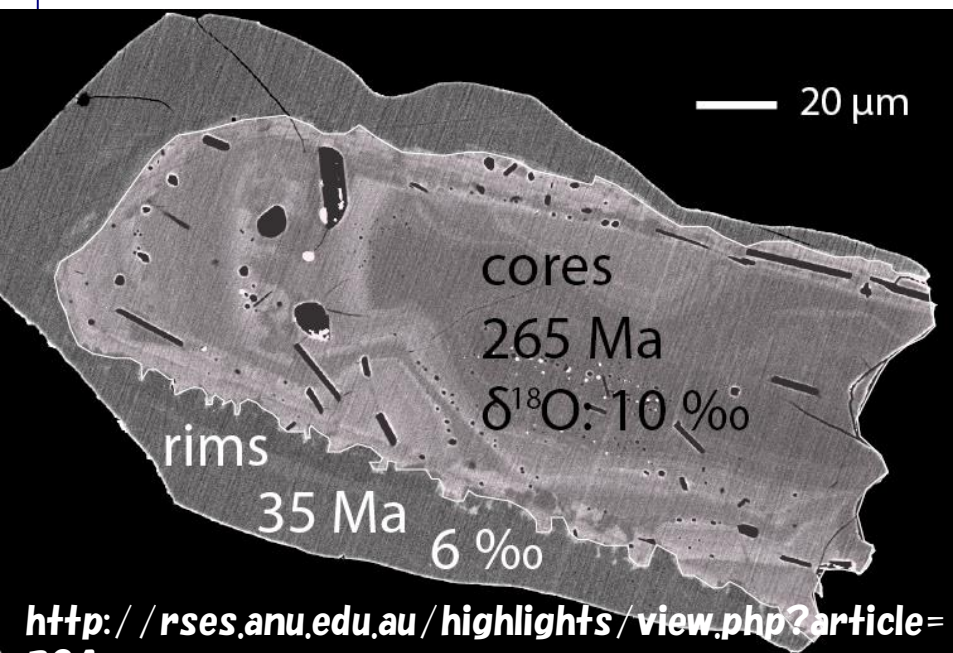
The SHRIMP-RG is at Stanford University as a result of a partnership between the U.S. Geological Survey and Stanford School of Earth, Energy & Environmental Sciences. The laboratory has been jointly operational since 1998, supporting scientists and students from the USGS, Stanford, and external visitors from around the world who





Photos by Prof. Fujioka

SHRIMP II at ANU (Australian National University) Geoscience Lab.



SHRIMP II SENSITIVE HIGH RESOLUTION ION MICROPROBE

SHRIMP II Schematic Diagram

The microbeam is the second component. It filters secondary ions according to their mass. Fold strength is varied in order to transmit a particular mass through to the collector.

The acceleration system (SSA) is the new component of the sensitive high resolution mass spectrometer. It filters secondary ions according to their energy.

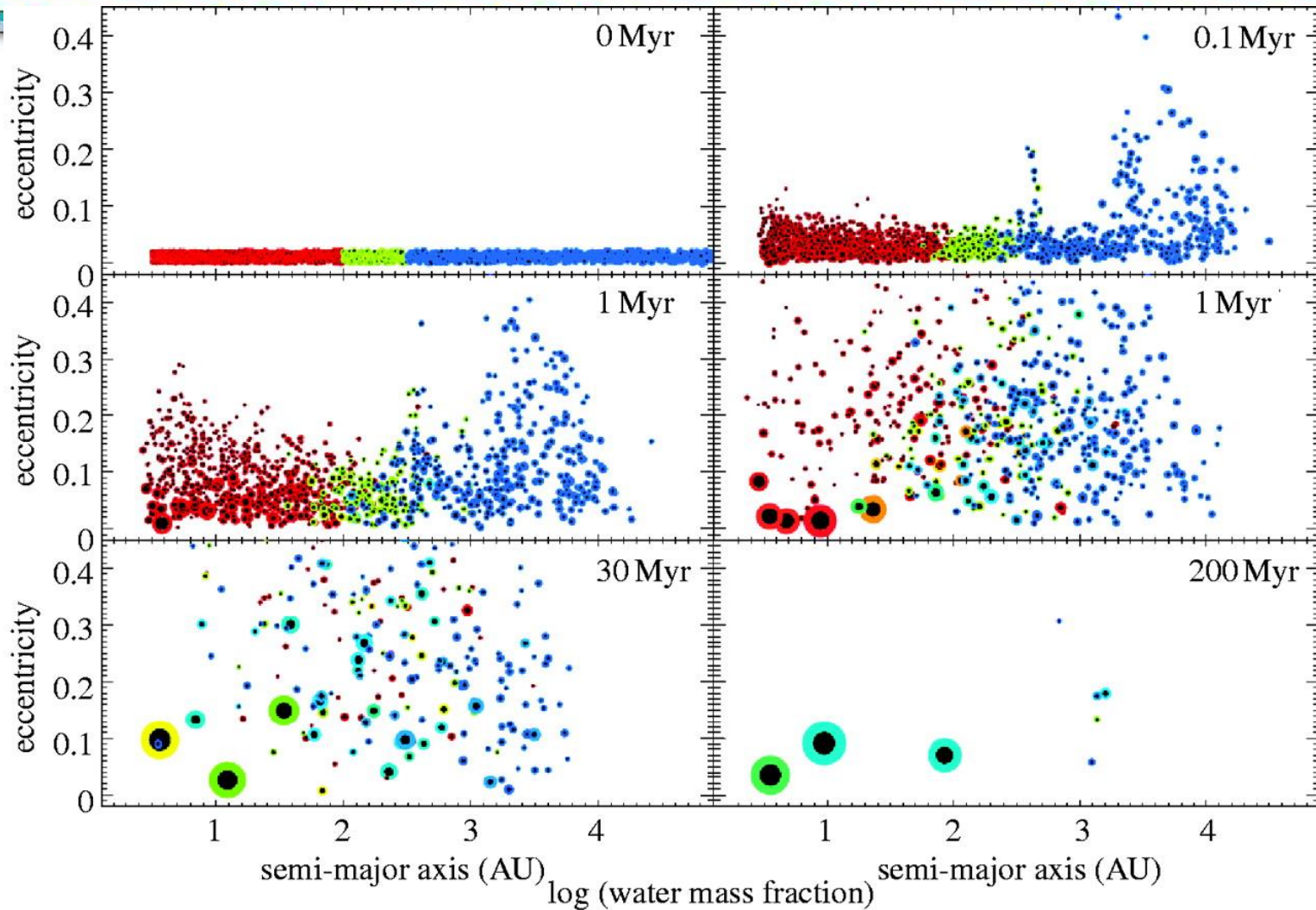
The ions of the sample are ionized in the primary column. The primary ions are then accelerated through 10000 volts into the secondary column.

The acceleration to the ions of primary ions which are accelerated through 10000 volts into the primary column.

The primary column consists of electrostatic deflectors and lenses to focus the primary ion beam to a diameter of approximately 20-30 microns on the sample.

The collector counts the number of secondary ions of a particular mass (e.g. Pb206 or U238) for a specified trace (e.g. Th and Sr).

Example of the formation of terrestrial planets from a series of asteroidal-to-Mars-sized bodies using a symplector integrator simulation as described in the text.

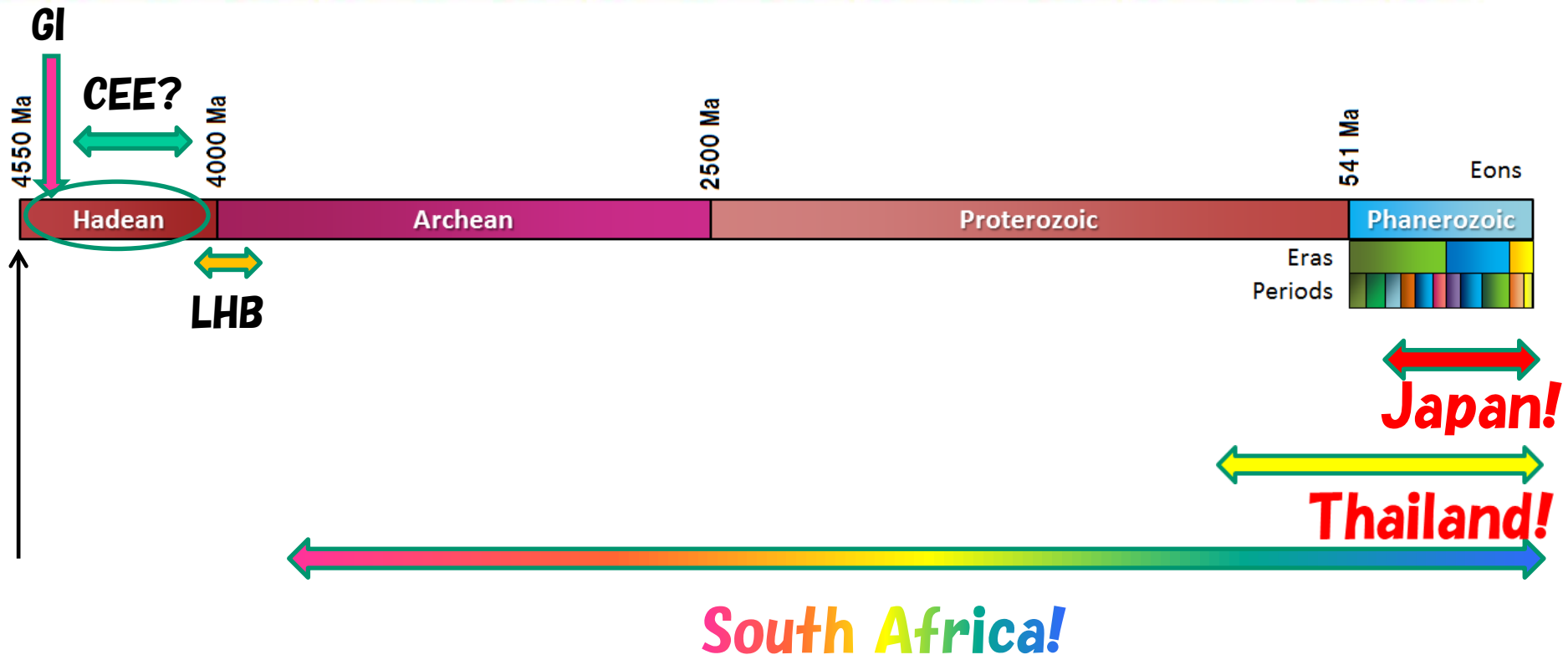


Example of the formation of terrestrial planets from a series of asteroidal-to-Mars-sized bodies using a symplector integrator simulation as described in the text. Six snapshots in time show the eccentricity versus semi-major axis of 1885 objects that collide, coalesce and grow under the perturbing influence of Jupiter and Saturn and their own growing gravitational fields. The size of each body is proportional to its linear diameter as it grows, and the amount of iron it contains is shown in black. Colours allow the eye to track mixing across regions of the solar system, and can also be considered a rough indication of the amount of water assumed present in these objects at the beginning, and in the bodies as they collide and grow, with the water mass fraction scale shown at the bottom. The final system is not precisely our own terrestrial planet system, but similar, and outcomes vary dramatically as boundary conditions are changed. From Raymond et al. (2006).

Jonathan Lunine *Phil Trans R Soc B* 2006; 361: 1721-1731

© 2006 The Royal Society
<http://rstb.royalsocietypublishing.org/content/361/1474/1721>

Early Earth (Part 1) Hadean eon



Forming of Earth 4.6Ga

Giant Impact (the birth of Moon) 4.5Ga

Cool Early Earth 4.4-4.0 Ga

Late Heavy Bombardment (LHB) 3.9 Ga



W.K.Hartmann (right) and Donald R. Davis(left)
as a founder of “Giant Impact Hypothesis”; the most plausible
theory of the origin of the moon.





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Giant impact hypothesis

From Wikipedia, the free encyclopedia

"Big splash" redirects here. For other uses, see [Big Splash \(disambiguation\)](#).

The **giant impact hypothesis** states that the [Moon](#) was formed out of the debris left over from a collision between the [Earth](#) and a body the size of [Mars](#), approximately 4.5 *Gya* (four and a half billion years ago). The colliding body is sometimes called **Theia**, for the mythical [Greek Titan](#) who was the mother of [Selene](#), the goddess of the Moon.^{[1][2]}

The giant impact hypothesis is the currently-favoured [scientific hypothesis](#) for the formation of the Moon.^[3] Supporting evidence includes: the Earth's spin and Moon's orbit having similar orientations,^[4] Moon samples indicating the surface of the Moon was once molten, the Moon's relatively small [iron core](#), lower density compared to the Earth, evidence of similar collisions in other star systems (that result in [debris disks](#)), and that giant collisions are consistent with the leading theories of the [formation of the solar system](#). Finally, the stable isotope ratios of lunar and terrestrial rock are identical, implying a common origin.^[5]

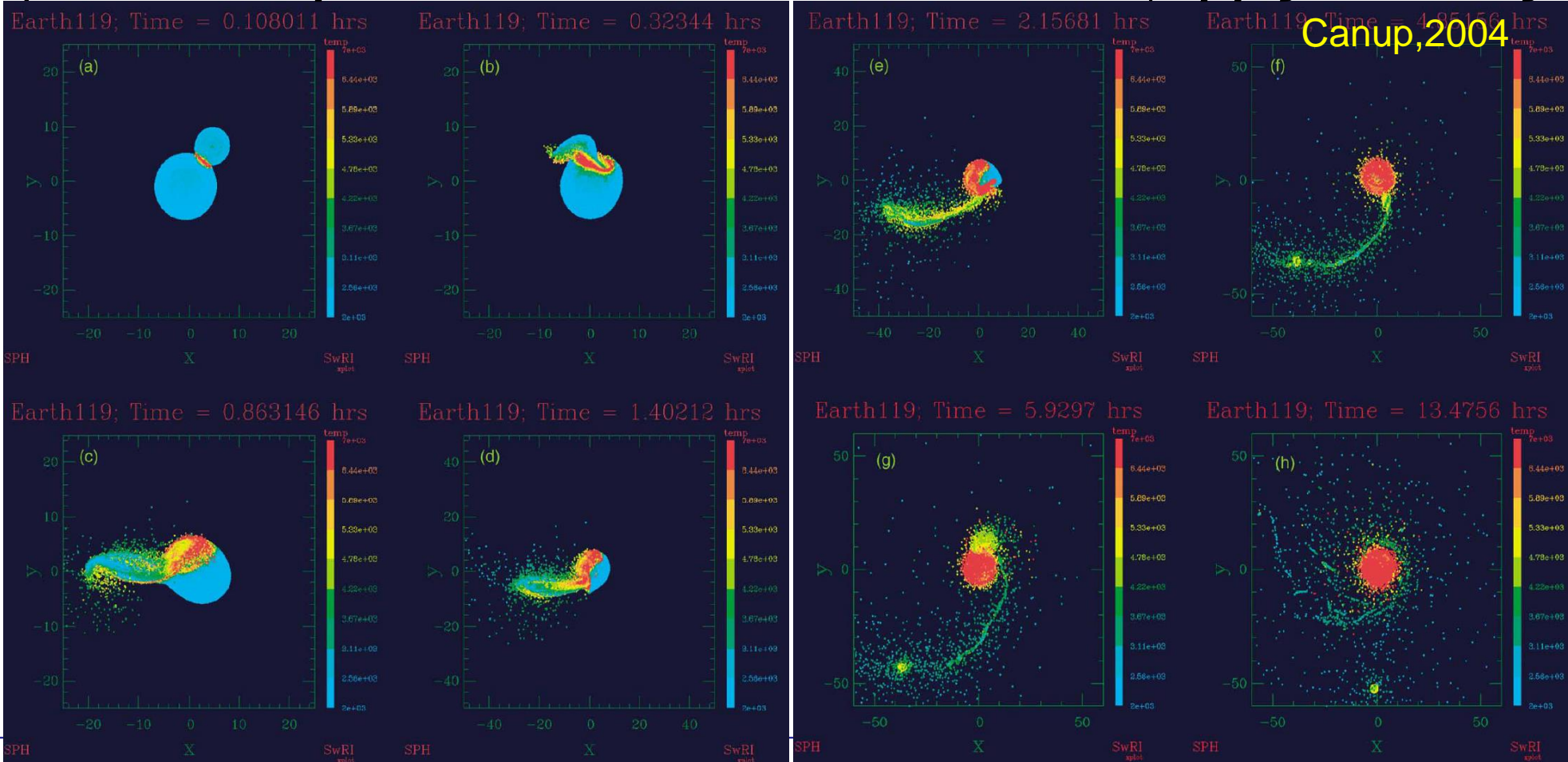
There remain several questions concerning the best current models of the giant impact hypothesis, however. The energy of such a giant impact is predicted to heat Earth to produce a global 'ocean' of [magma](#); yet there is no evidence of the resultant [planetary differentiation](#) of the heavier material sinking into Earth's mantle. At present, there is no self-consistent model that starts with the giant impact event and follows the evolution of the debris into a single moon. Other remaining questions include when the Moon lost its share of [volatile elements](#) and why [Venus](#), which also experienced giant impacts during its formation, does not host a similar moon.



Artist's depiction of a collision between two planetary bodies. Such an impact between the Earth and a Mars-sized object likely formed the Moon.

Supporting evidence of GI (wiki)

- i) Earth's spin and the Moon's orbit have similar orientations.
- ii) Moon samples indicate that the Moon once had a molten surface.
- iii) The Moon has a relatively small iron core.
- iv) The Moon has a lower density than Earth.
- v) Evidence exists of similar collisions in other star systems (that result in debris disks).
- vi) Giant collisions are consistent with the leading theories of the formation of the solar system.
- vii) The stable-isotope ratios of lunar and terrestrial rock are identical, implying a common origin





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Giant impact hypothesis

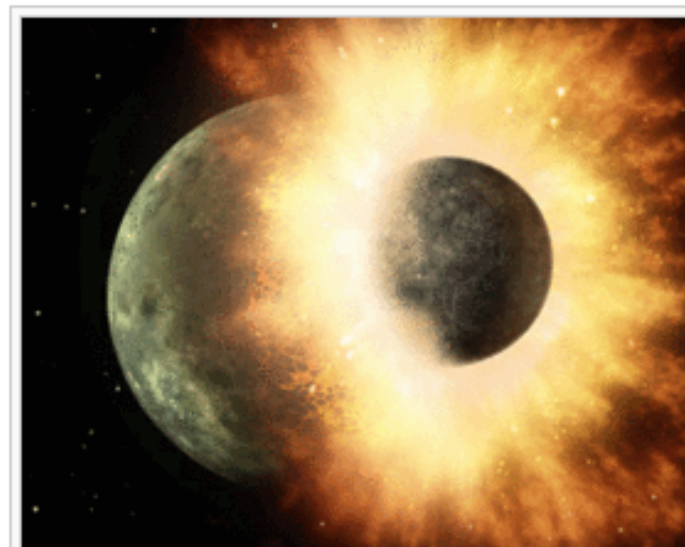
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Artist's depiction of a collision between two planetary bodies. Such an impact between the Earth and a Mars-sized object likely formed the Moon.

Celebrating 40 Years of Scientific Discovery



米国惑星科学研究所創始者
W.K.Hartmann氏



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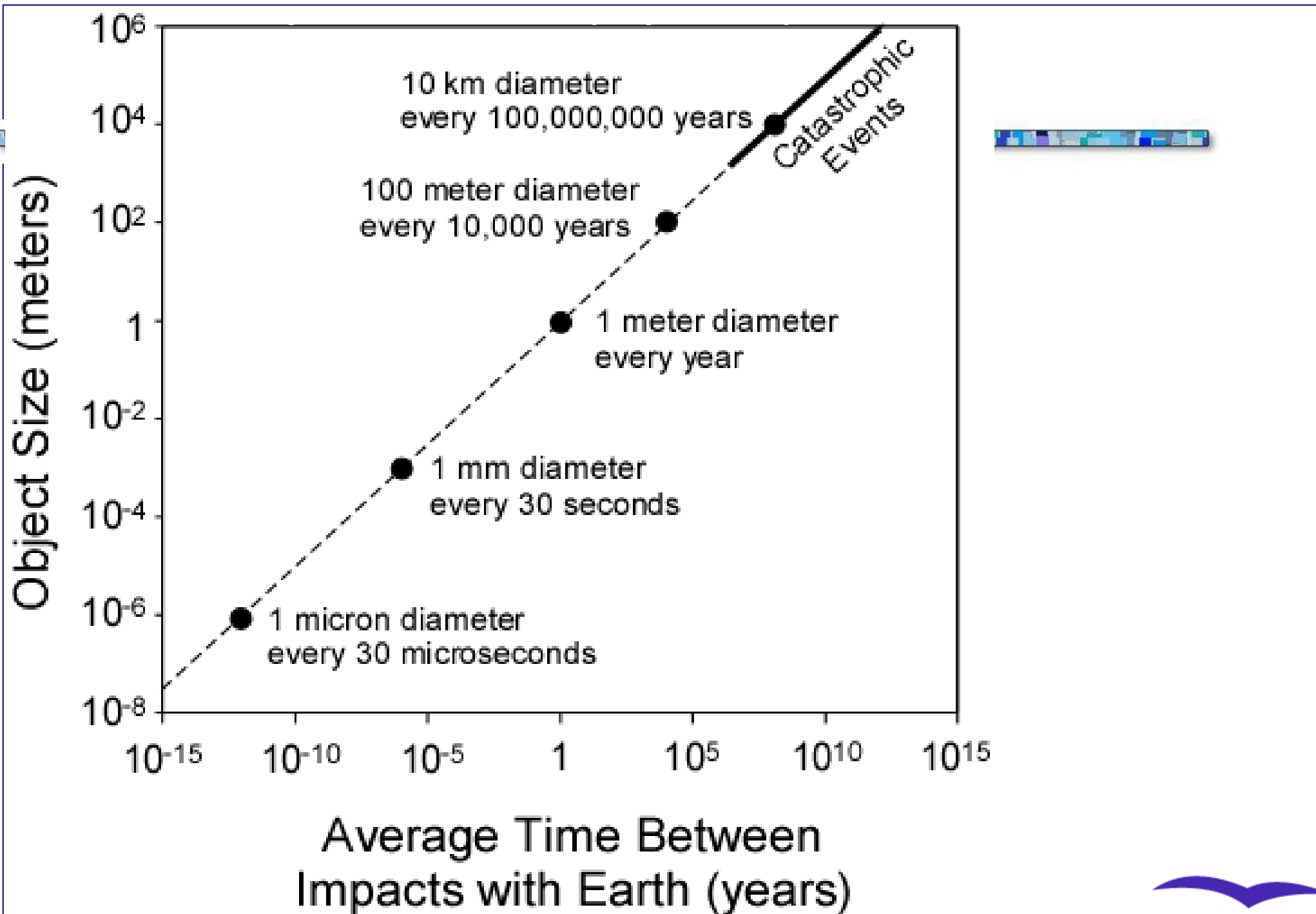
Planetary
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Institute
Est. 1972











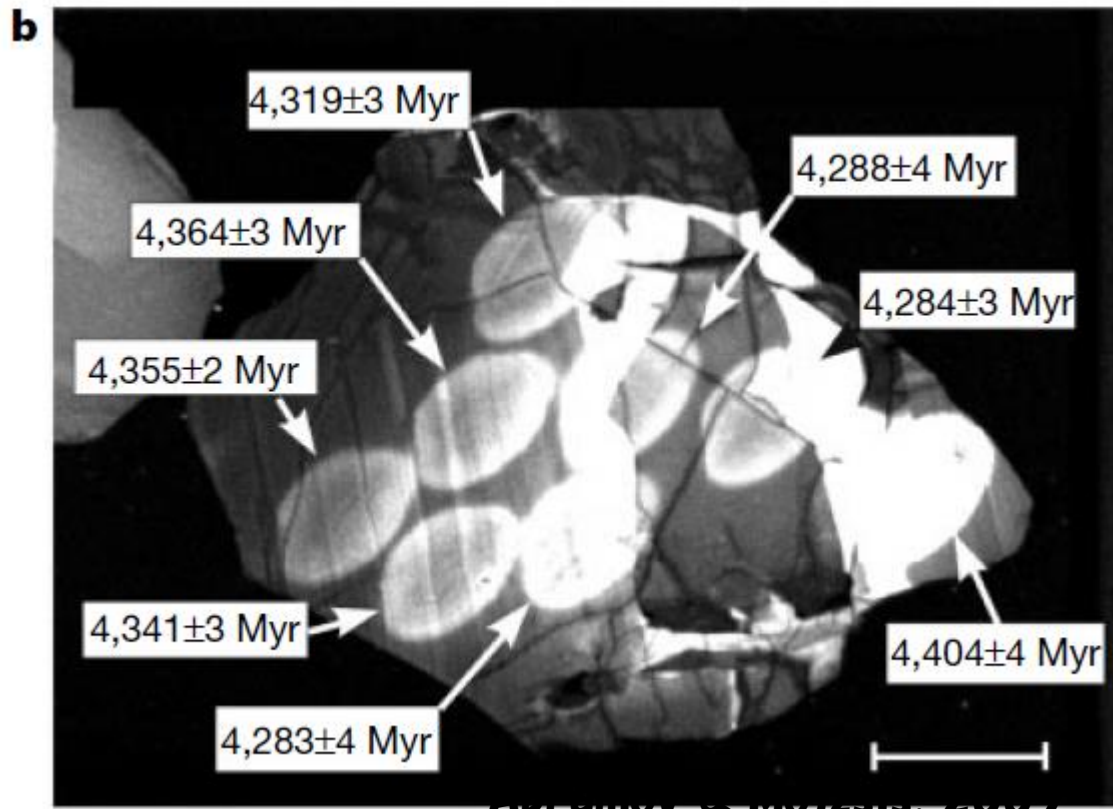
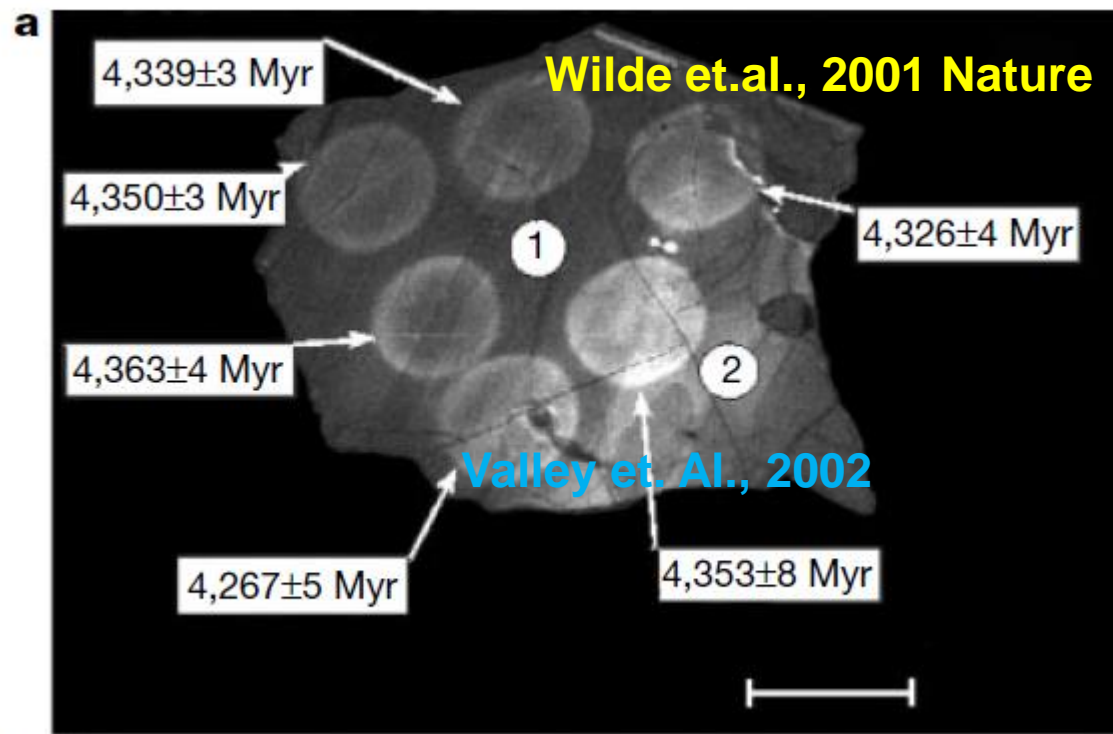
4.4Ga Zircon by SHRIMP II

**Oldest mineral
Granitic rocks**

Why Zircon?

**Resist against
weathering
U, Pb rich**

Lineweaver & Norman, 2008



ADAMOV & MOJZIS, 2007

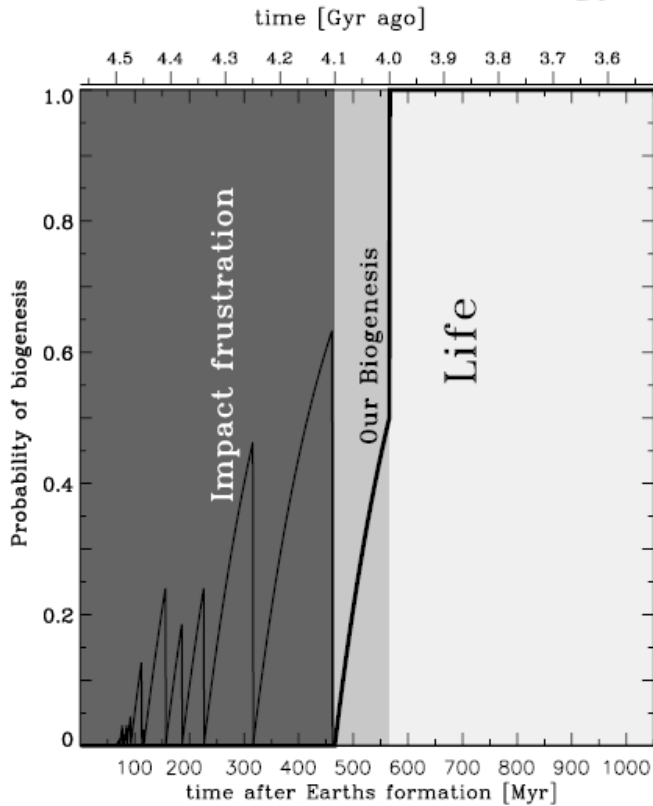
A cool early Earth

John W. Valley*
William H. Peck*
Elizabeth M. King*

Department of Geology and Geophysics, University of Wisconsin, Madison, Wisconsin 53706, USA
Simon A. Wilde

School of Applied Geology, Curtin University of Technology, GPO Box U1987, Perth, Australia

Sterilizing Impacts and life "Panspermia Hypothesis"



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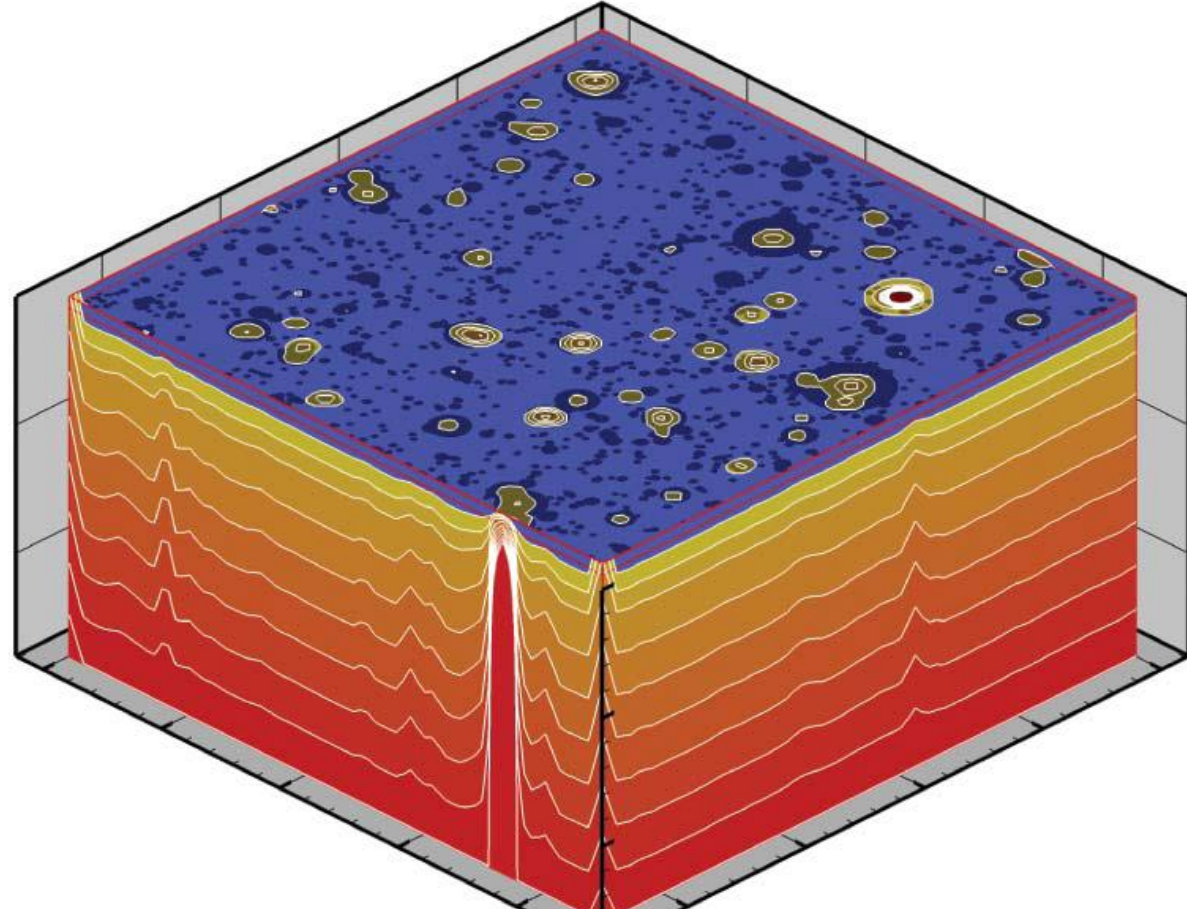


Fig. 2. After the sterilizing impact that formed the Moon about 90 ± 20 Myr after the formation of the solar system (Halliday 2008), a heavy but decreasing and stochastic bombardment lasted for a few hundred million years probably frustrated the origin of life on Earth. Eventually, the molecular evolution that led to life as we know it, was able to squeeze through the thermal bottlenecks produced by impacts (however see Abraomov & Mojzsis 2008a,b). Figure from Davies & Lineweaver 2005.

Hadean Earth (4.0 Ga)

Hadean Earth

ca. 4 billion years ago



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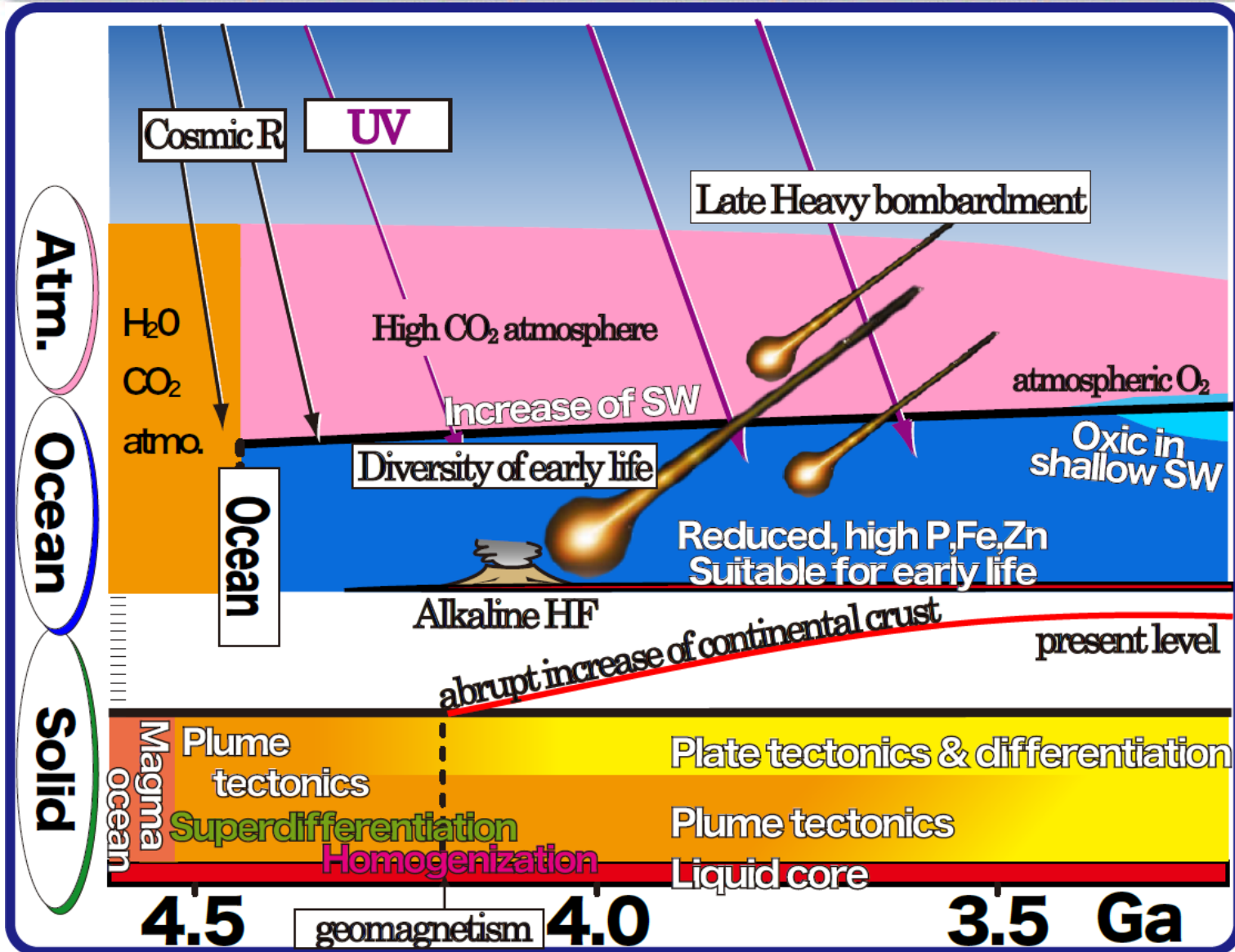


Dr. Simone Marchi kindly allow me to use this gif-images.

<http://www.boulder.swri.edu/~marchi/>

Figure 1. An artistic conception of the early Earth-Moon system. The Earth is pictured as surface pummeled by large impacts, resulting in extrusion of impact-generated deep-seated magma onto the surface. At the same time, distal portion of the surface could have retained liquid water. The Moon is pictured as a dry, heavily cratered body. The Moon is far less geologically active than the Earth and its older surface and rocks have been used to calibrate our bombardment.



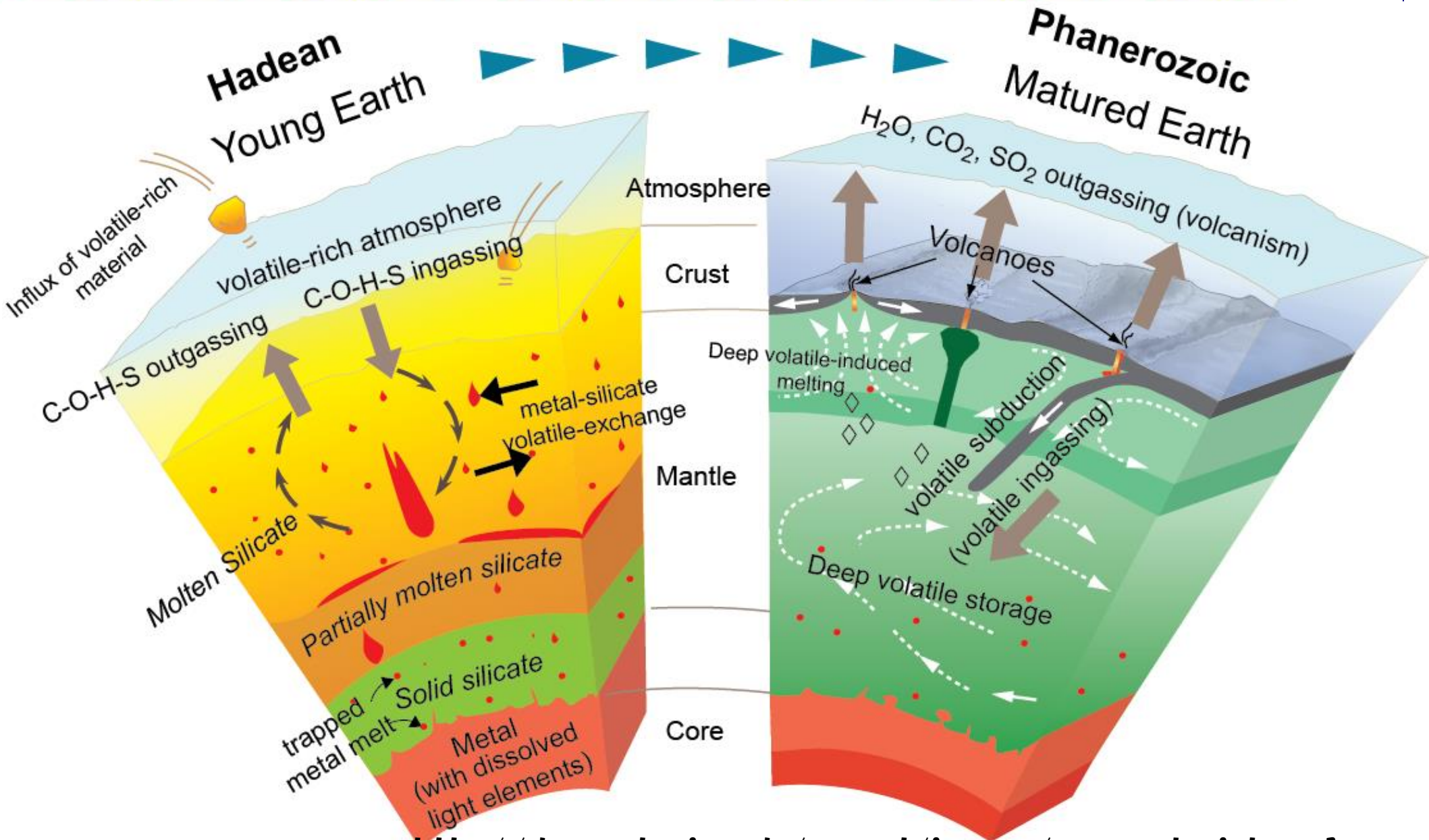


**A. Prof. Komiya
(Tokyo Univ.)
kindly provides
me this image.**

https://www.jsps.go.jp/j-grantsinaid/12_kiban/ichiran_26/j-data/h26_j3213_komiya.pdf



Formation and evolution of earth



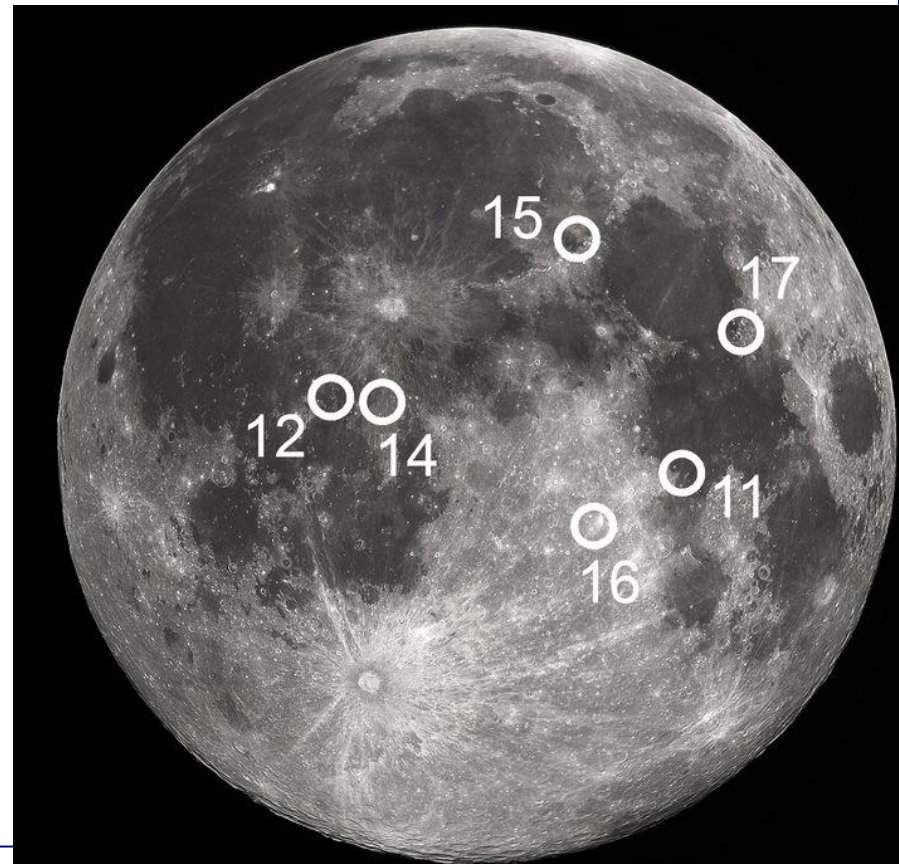
Late Heavy Bombardment (LHB)

<Evidence>

Moonrocks: Apollo mission

The ages of **impact melts** collected at **these sites clustered between about 3.8 and 4.1 Ga**. The apparent clustering of ages of these led to postulation that the ages record an intense bombardment of the Moon. They called it the 'lunar cataclysm' and proposed that it represented a dramatic increase in the rate of bombardment of the Moon around **3.9 Ga**.

http://public.media.smithsonianmag.com/legacy_blog/age-histogram.jpg



“クレータ年代学”

Crater Chronology: principle

- **Basic idea by W. K. Hartmann in 1960's**
- **The principle is quite simple!**
 - > Heavy cratering surface is old.

Lunar Reconnaissance Orbiter Wide-Angle Camera Mosaic

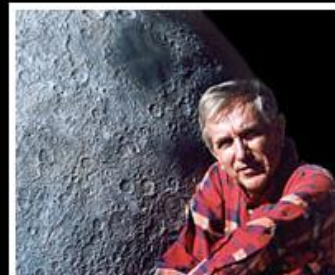
Few Craters = Younger

Many Craters = Older

Relative Age: Something is younger or older than something else.

Stratigraphy: Geologically above or below something

Absolute Age: Assigning an actual age, like 300 million years



*Bill Hartmann's
Home Page*



<http://sservi.nasa.gov/articles/video-the-lunar-crater-chronology>

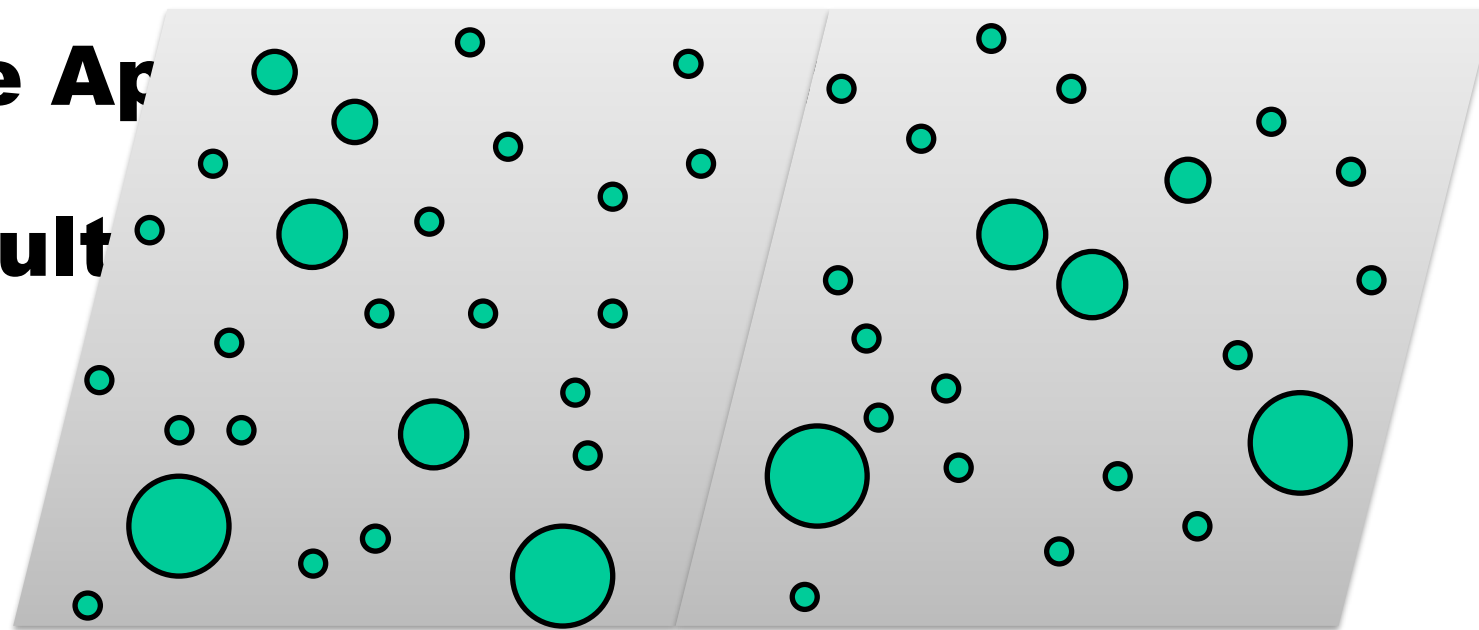


from 2011 SSH summer meeting presentation

What is Crater-Chronology?

William K. Hartmann (PSI) first developed isochrones diagram for the Moon.

➤ **The Ap
result**



➔ **Old**

➔ **Young**

Confirmation of “Crater counts”

After some simplified assumptions, he completed an Isochrones chart for the Moon in 1960's. From it, he estimated the surface age of “Luna maria” as 3.6 Giga years. “月の海が36億年を示す”

Five years after, the Apollo mission brought back many moon rock samples and confirmed the reliability of this method by measuring the radiometric ages of these rocks.

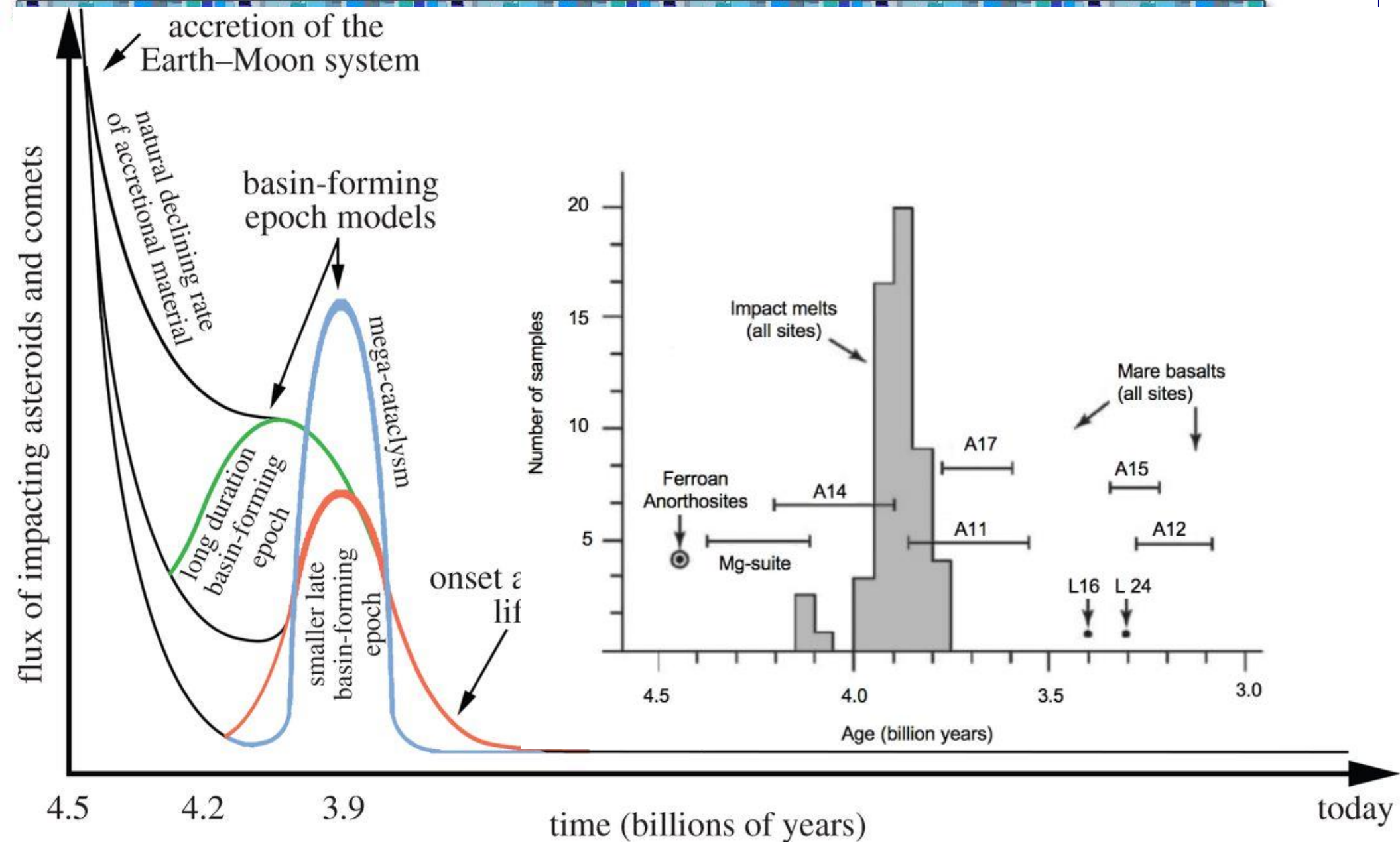


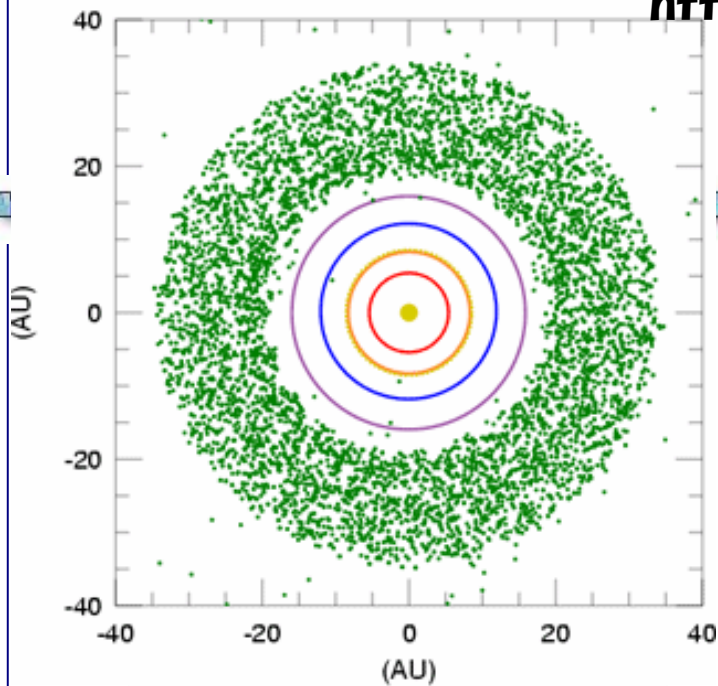
Google **Q1. Another feature of Craters?** © 2013 Google. All rights reserved. This view

Crater size distribution?



Late Heavy Bombardment part 2.





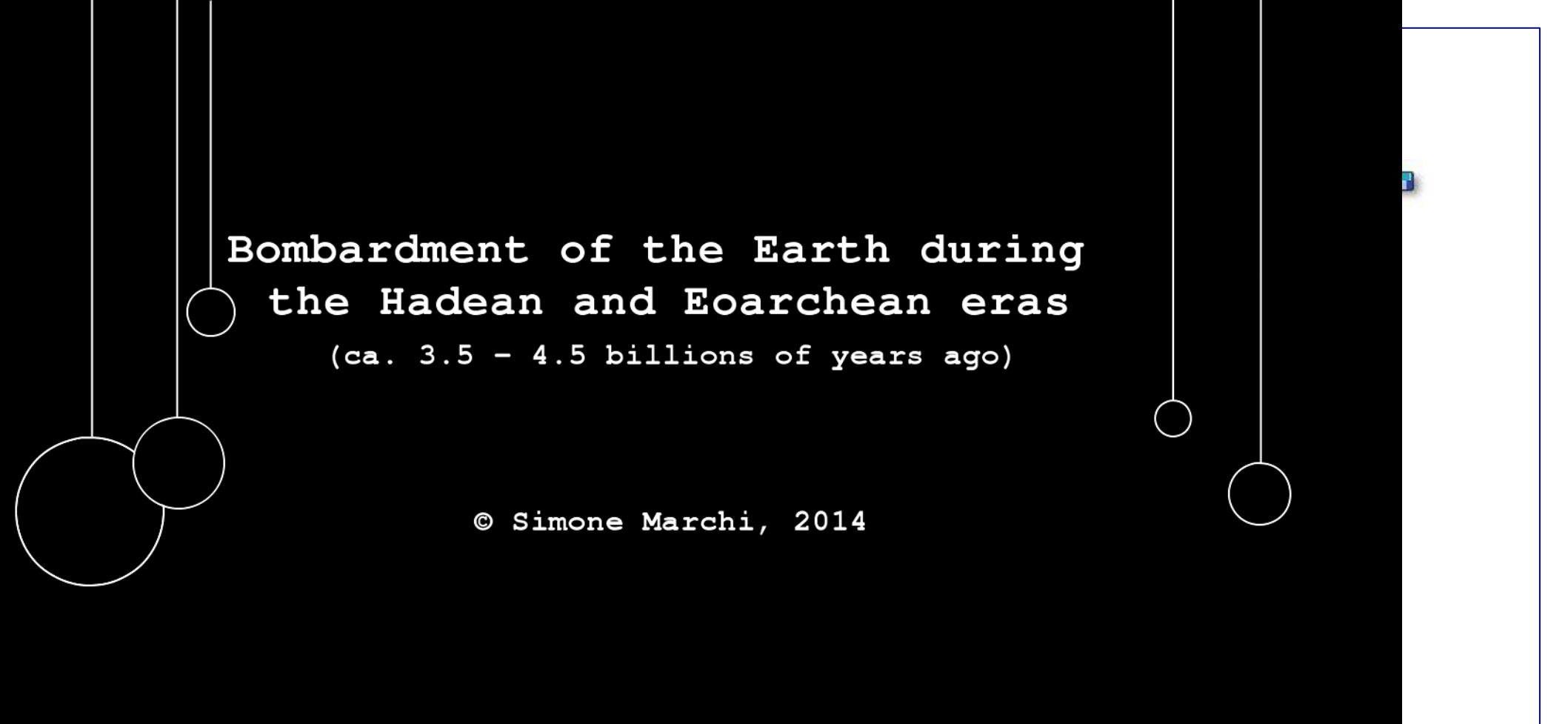
(From Gomes, et al., 2005, *Nature*, v. 435, p. 466-469.)

Simulation : “Nice model” R. Gomes et.al., *Nature* 2005 A migration of the giant planets

In this dynamical simulation of the late heavy bombardment, the Sun is in the center, the colored circular rings represent the orbits of the four giant planets, and the green dots represent the disk of planetesimals between 15.5 AU and 34 AU.

Each panel represents the state of the planetary system at a different time, starting at $t=100$ million years. Saturn and Jupiter migrate slowly, reaching 2:1 resonance. This scatters Neptune and Uranus. Their extreme migrations scatter planetesimals in a short time interval – a cataclysm.

The four panels below correspond to four different snapshots taken from the simulations. From left to right: The beginning of planetary migration (100 Myr), just before the beginning of the scattering (879 Myr), just after scattering has started (882 Myr), and 200 Myr later, when only 3% of the initial mass of the disk is left and the planets have achieved their final orbits.



Bombardment of the Earth during the Hadean and Eoarchean eras

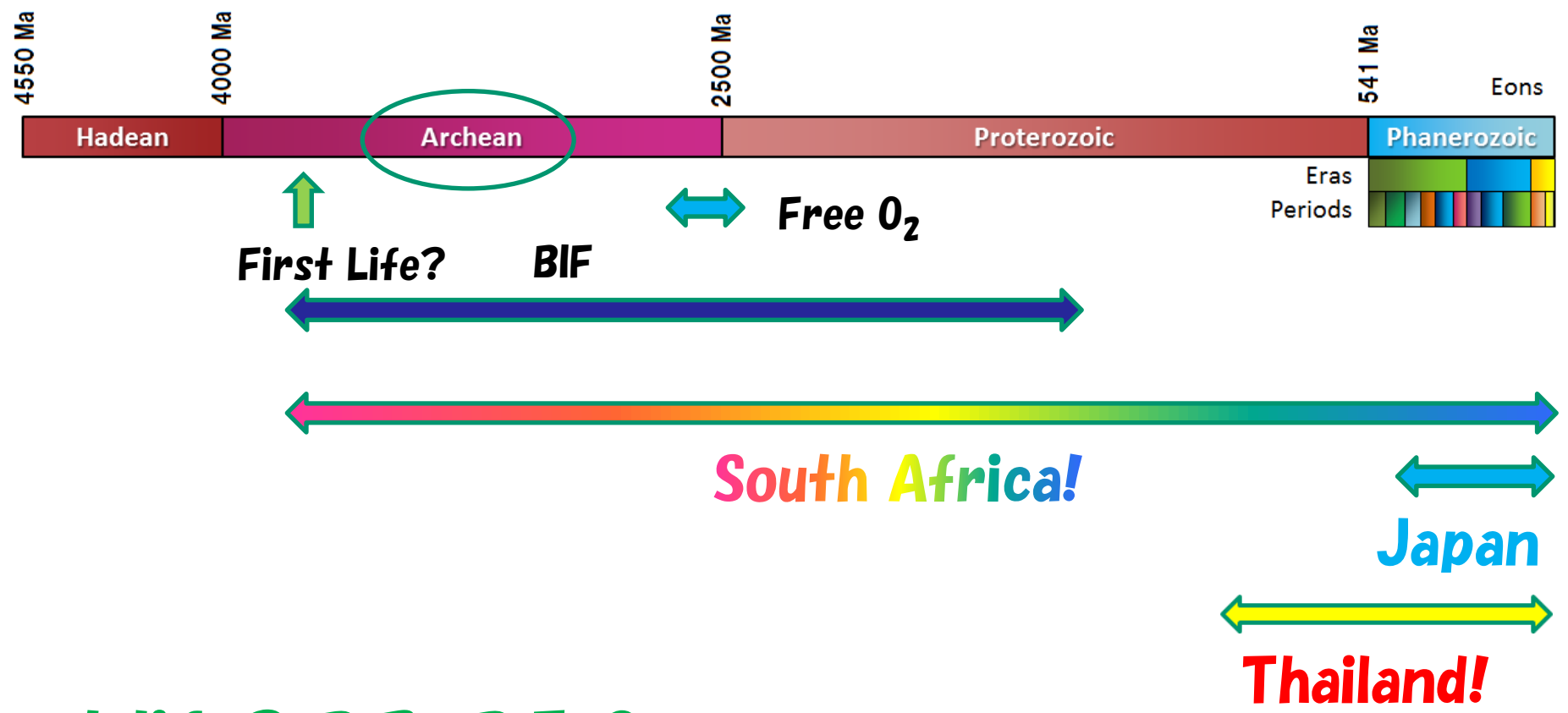
(ca. 3.5 - 4.5 billions of years ago)

© Simone Marchi, 2014

Figure 2. An animation showing the effects of bombardment on the early Earth. Each circle represents the area highly processed by an impact. The diameters of the circles correspond to the final size of the craters for impactors smaller than 100 km in diameter, while for larger impactors it corresponds to the size of the region buried by impact-generated melt, as described in the text. Color coding indicates the timing of the impacts. The smallest impactors considered have a diameter of 15 km.

<http://www.boulder.swri.edu/~marchi/>

Early Earth (Part 2) Archean eon



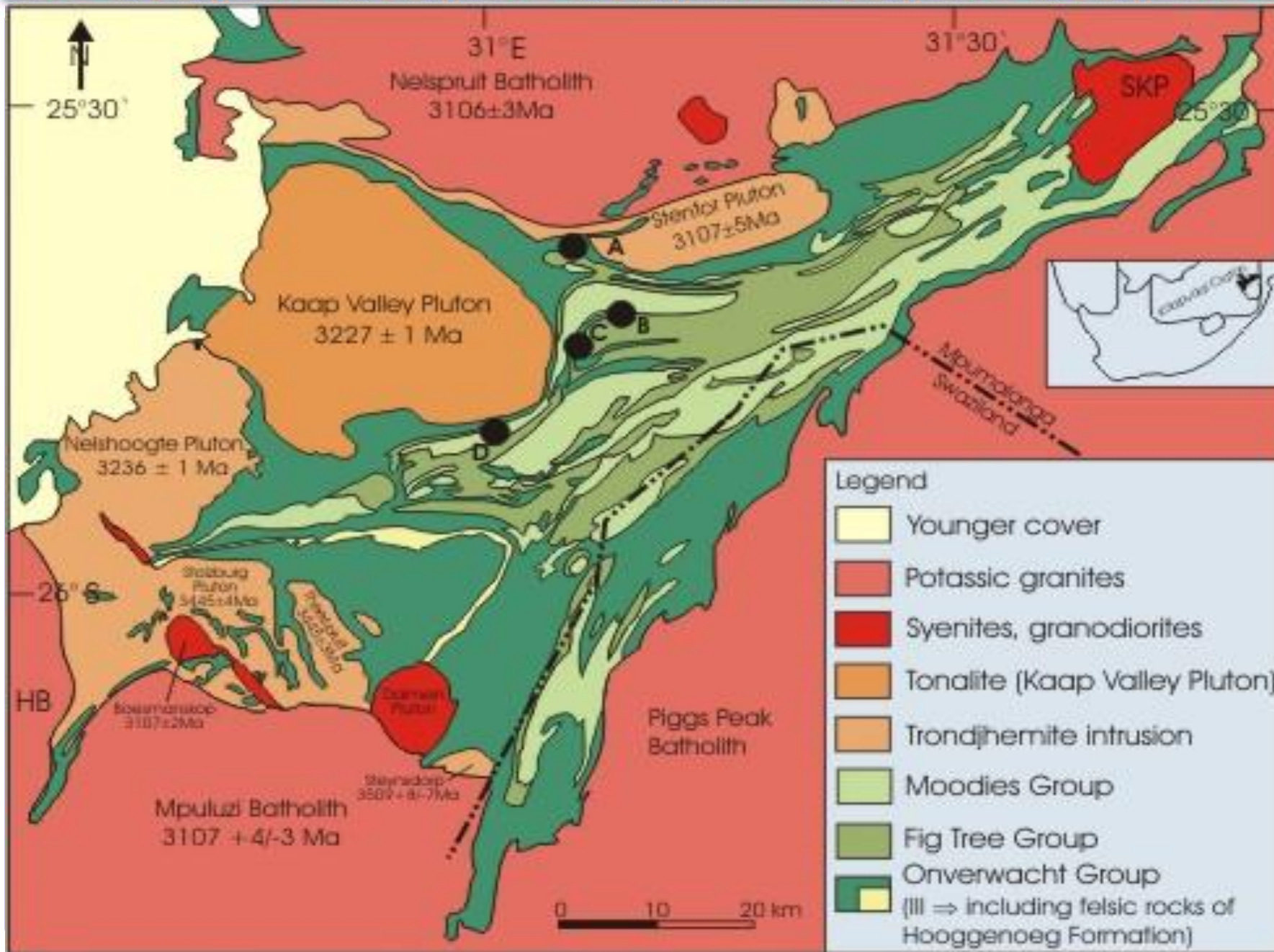
First life? 3.7-3.5 Ga

Banded Iron Formation (BIF) 3.8 to 1.9 Ga

Free Oxygen 2.5-2.4 Ga



Barberton Geological Map



Barberton Field Trip (2010)

Oldest Craton (> 3.6Ga)

Komatiite (Mg20% Ultra Mafic, 3.5Ga)

1600°C (High Temp. Magma) → Spinifex Texture

Pillow Lava : evidence of Oldest Oceanic Crust

**Oldest Sedimentary Rocks bearing no metamorphism
(3.4–3.2Ga)**

Sand–Mud Tidal Rhythmite

Chart (BIF)

Bio mats (Javaux, Nature 2010)

Spherules of A large Impact (>30km)

Oldest Gold Ore

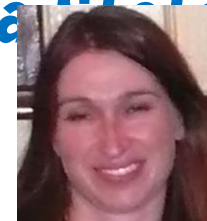
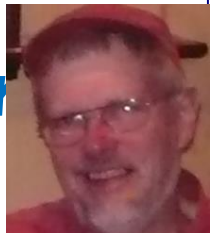
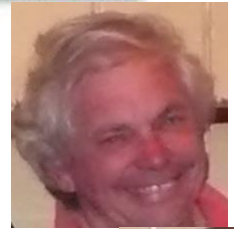


Members



Members

- **Dexter Perkins** : Professor, Dept. Of Geology and Geological Engineering, Univ. of North Dakota
- **Cathy Manduca** : Director, Science Education Resource Center – Carleton College
- **Karl Wirth** : Associate Professor, Geology Department, Macalester College St. Paul, Minnesota
- **Donald P. Schwert** : Professor of Geology, Center for Science & Mathematics Education, North Dakota State University
- **Florence le Hebel** : Université Lyon 2
- **Jim Nicholls, Retired** : Department of Geology and Geophysics, University of Calgary
- **Dr. Dion Brandt** : Geological Consultant (Driver and leader)
- **And Me** (USA 4, Canada 1, French 1, SA 1, Japan 1)



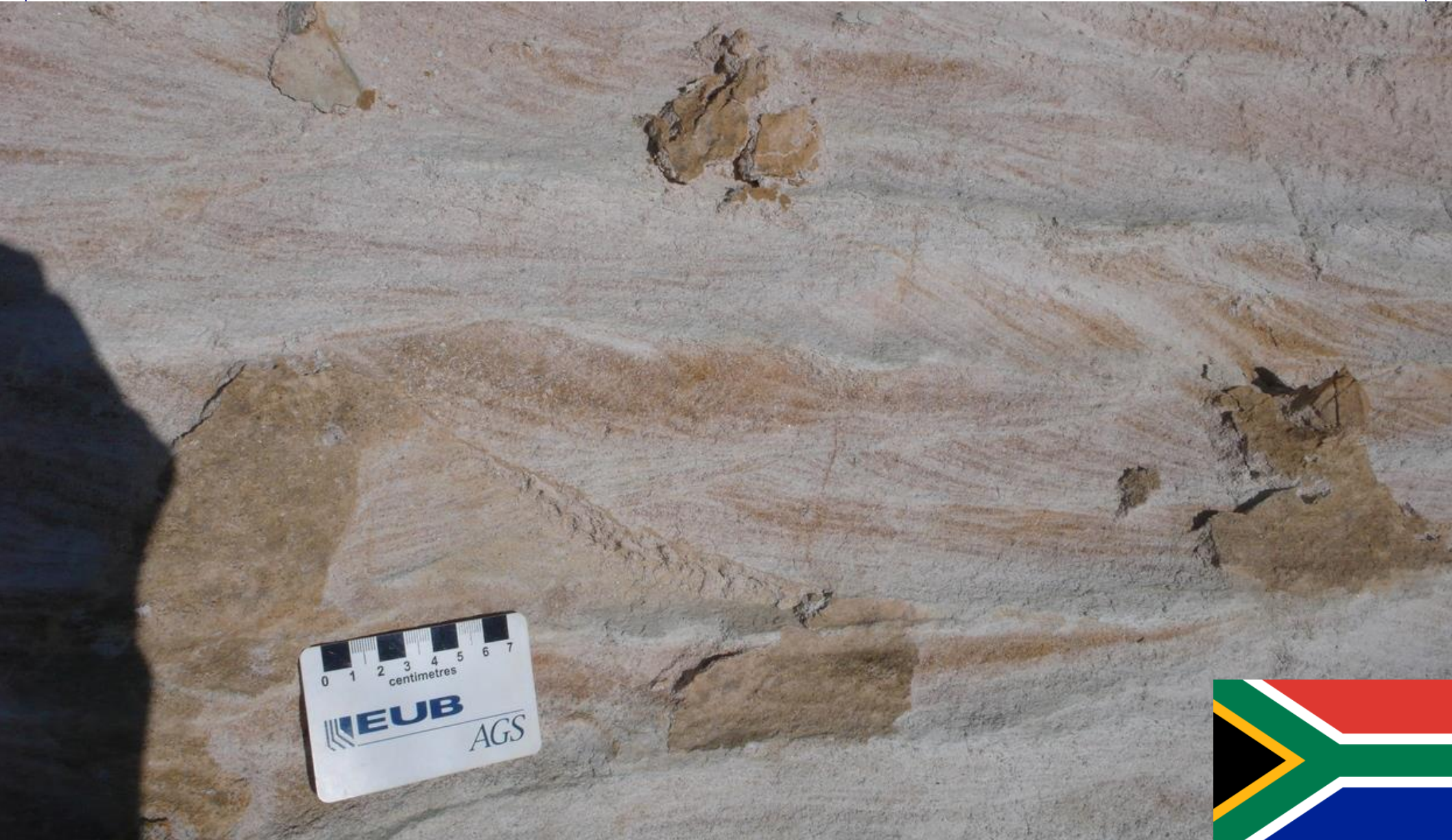
Barberton



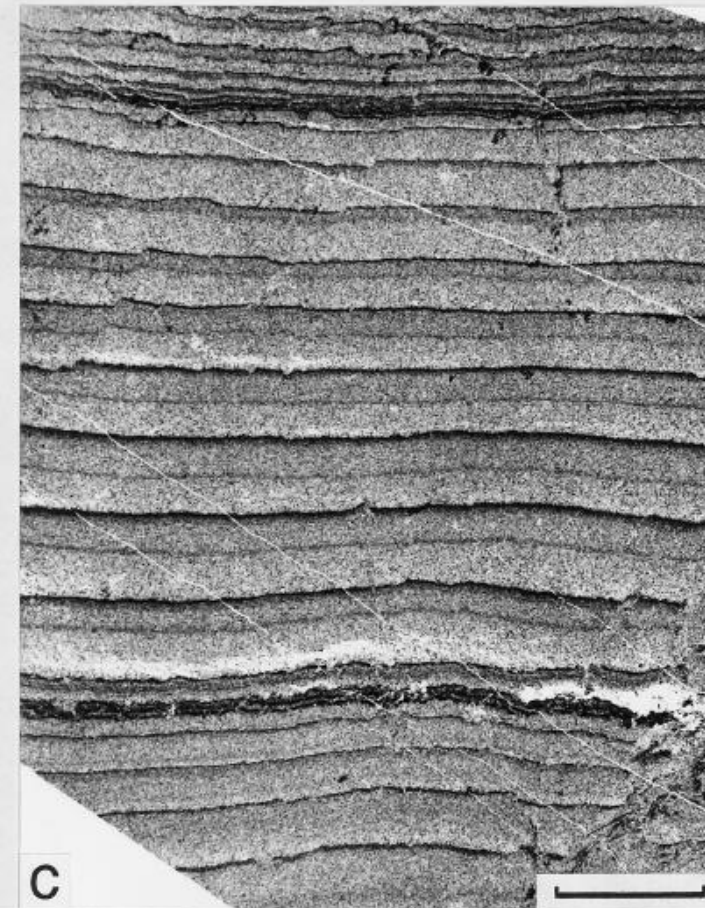
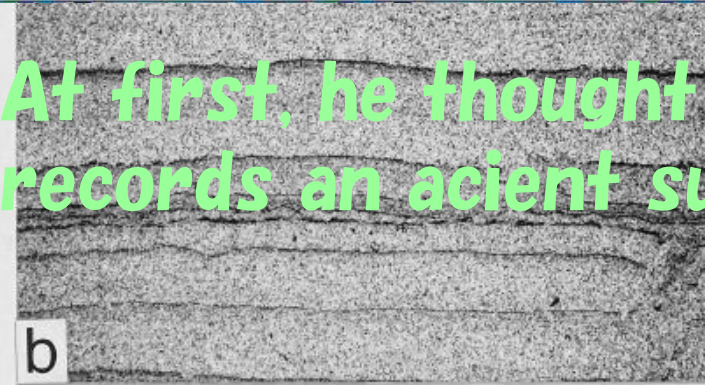
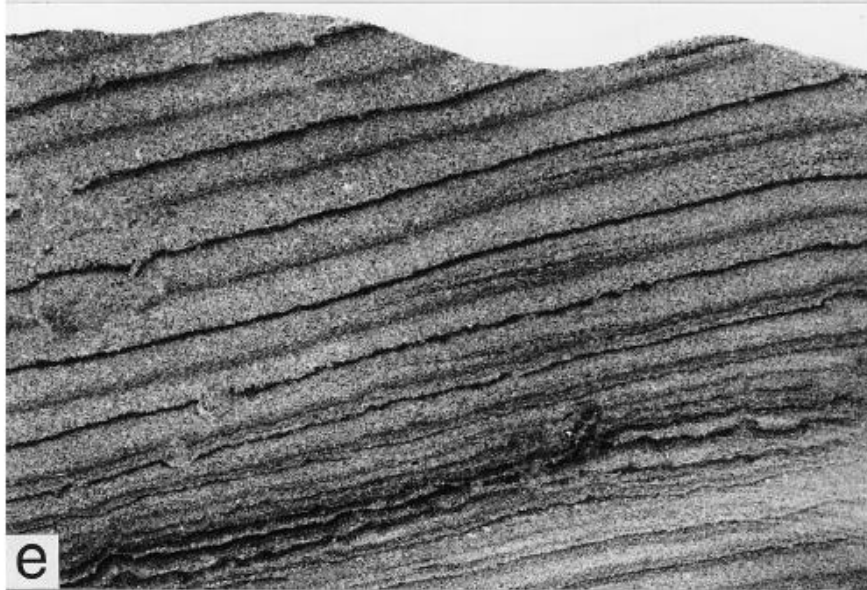
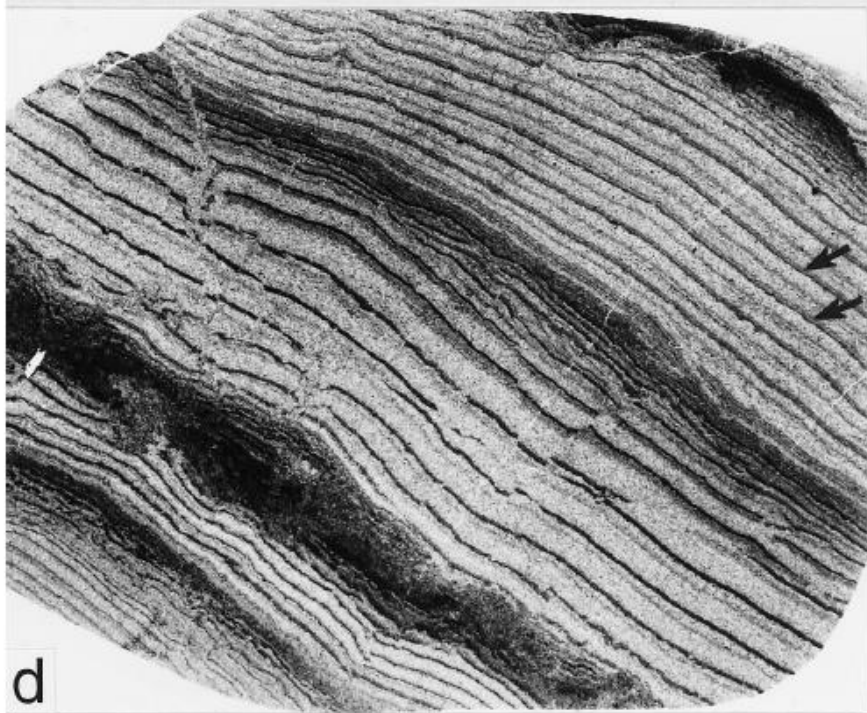
Moodies Group (3.2Gy) shallow marine tidal Rhythmites

South Africa 2010

At Barberton

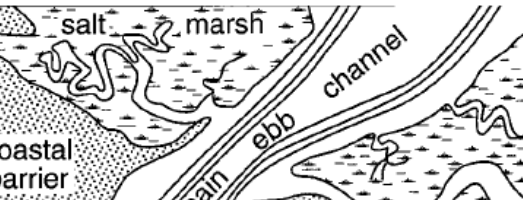
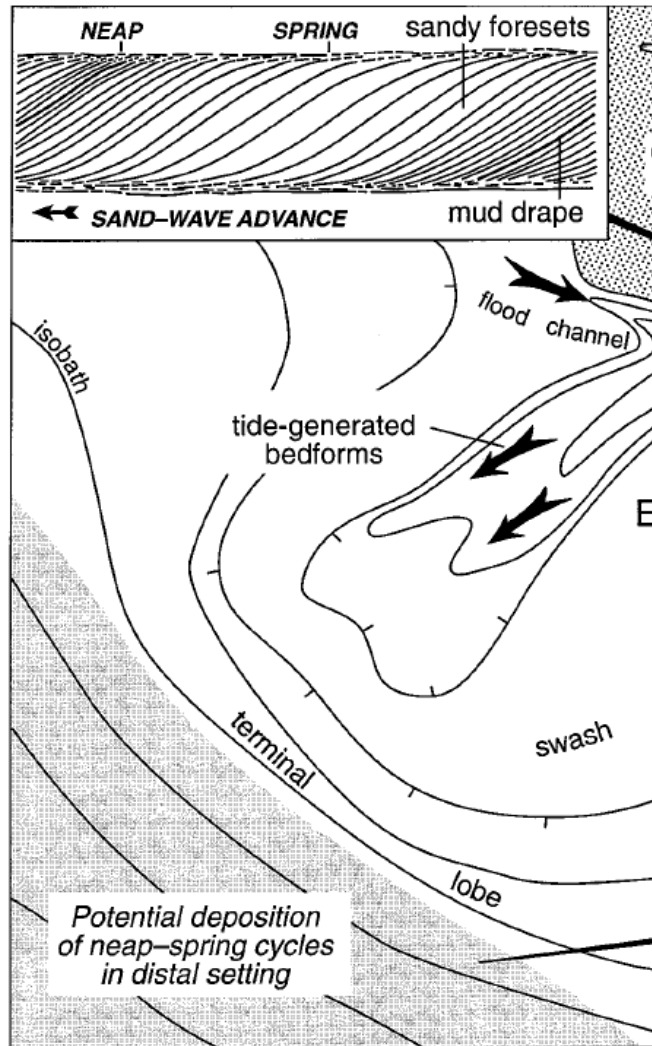


Williams 2000: Australia 0.6Ga



At first, he thought the lamina records an ancient sunspot cycle





48 • Williams: EARTH'S PRECAMBRIAN ROTATION

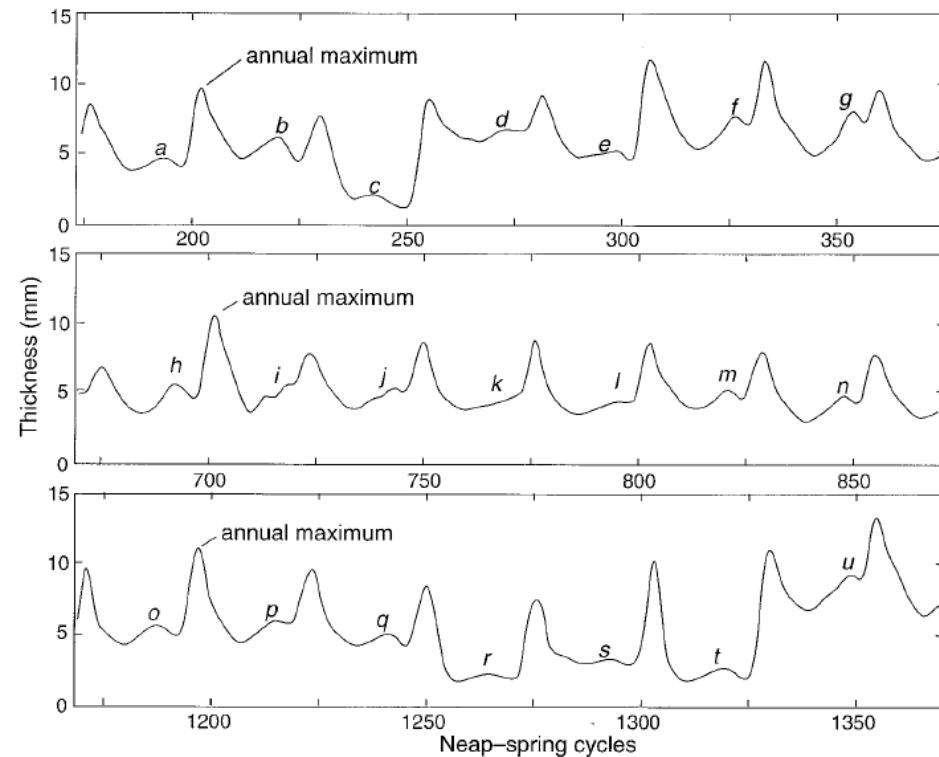


Figure 5. Envisaged environment of deposition for the E a hypothetical ebb tidal delta adapted from Imperato et al. [] tidal inlet, where fine-grained sediment is entrained by ebb t via the main ebb channel to deeper water offshore. There tl cycles of thin, graded laminae mostly of sand and silt (show the neap-spring cycles become progressively more abbreviat into marine shelf mud. Where protected from wave acti deposition and preservation of long rhythmite records. Ti inset) are confined to proximal, nearshore tidal channels. M the permission of the International Union of Geological S

Figure 9. Three extracts from the Elatina paleotidal record of neap-spring cycle thickness (smoothed by a five-point filter weighted 1, 4, 6, 4, 1; neap-spring cycle number increases up the stratigraphic succession), showing 24 first-order peaks that are equated with the nontidal annual or seasonal maximum in sea level. The plots span the three intervals where the second-order peaks (peaks a–u), which are interpreted as reflecting the semiannual paleotidal cycle, show minimal height (peaks c–e, j–l, and q–t); the symmetry of the annual peaks tends to be greatest at these places. Over the 60-year record, a period of 19.5 ± 0.5 years is revealed by variation in the height of the semiannual peaks (see Figure 13a), as measured from the base of the preceding trough to the top of the peak or to the midpoint between rare twin peaks (peaks i and j). Neap-spring cycle thickness shows a gradual decrease for cycles 1200–1325 and abrupt increases at cycles 250 and 1325; these nonperiodic changes evidently reflect sedimentary processes on the tidal delta such as a gradual blocking of the main ebb channel followed by channel avulsion.

Verify the Giant Impact Theory

**In the Archean eon,
Moon was more closer
to earth.**

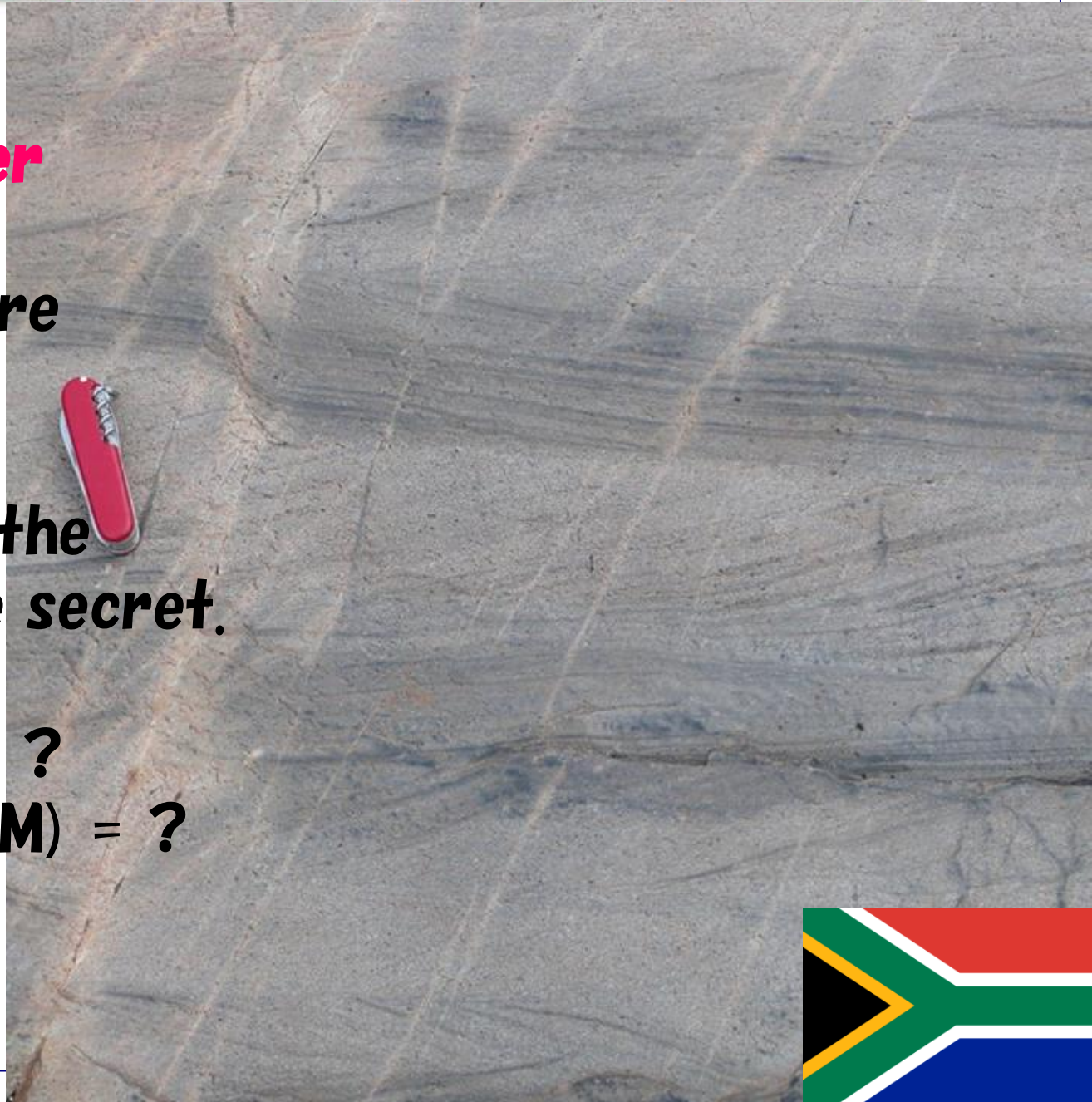
**So, these periods were
shorter than today.**

How to examine?

**The tidal records in the
sediments reveal the secret.**

length of day (LOD) = ?

Length of month (LOM) = ?



Moodies Group (3.2Gy) Analysis of tidal rhythmites

- 1) Rapid moon rotation: 18 days
- 2) Unknown: length of day and month
- 3) Circular orbit of Moon

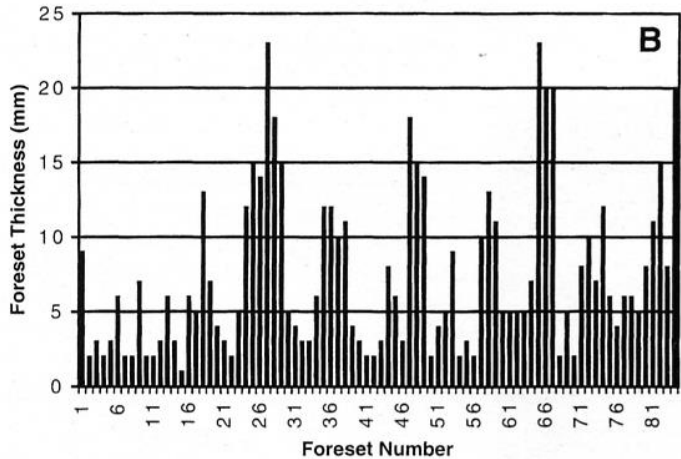
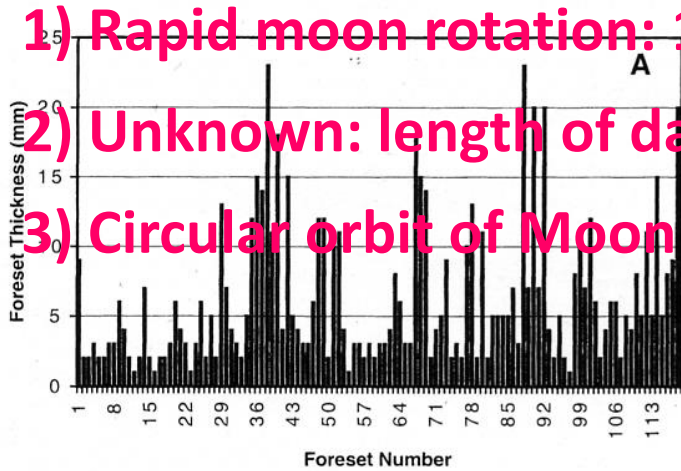


Figure 3. A: Traverse two—all data. Histogram of sandstone foreset bundle thicknesses plotted against foreset number for traverse two through cross-bed set shown in Figure 2. Note variation in thickness of sandstone foresets and common presence of thick-thin pairs of foresets. B: Traverse two—subordinates removed. Histogram of inferred dominant-tide foreset bundle thicknesses plotted against foreset number for traverse two through cross-bed set shown in Figure 2. Inferred subordinate flood-tide laminations were removed visually from data sets. Note that interpreted neap-spring-neap cycles are 9–10 days long and that alternate neap-spring-neap cycles are thicker and thinner, respectively.

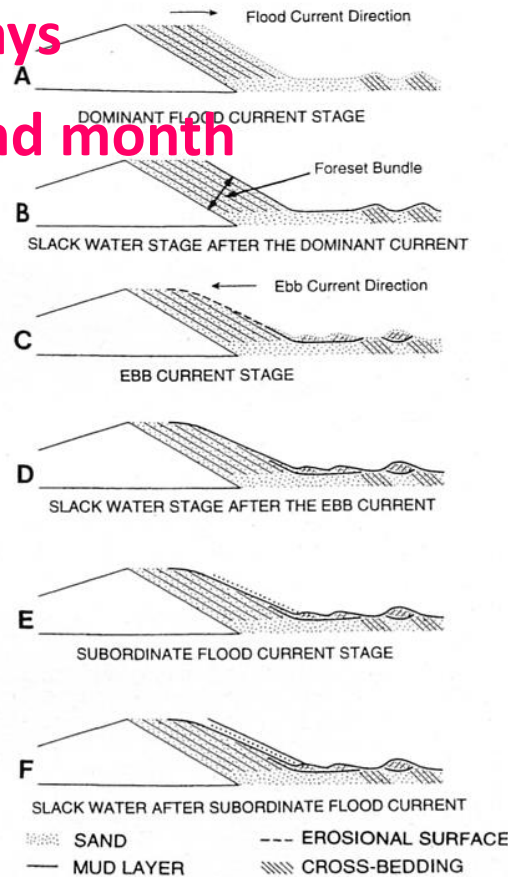


Figure 4. Migration of sand wave in tidal system characterized by strong flood current and weak ebb current (modified after Visser, 1980). Note that most sand is deposited on lee face of sand wave during dominant flood stage (A), whereas only thin sand layer is deposited on lee face during subordinate flood stage (E). Dominant and subordinate flood currents are typical of semidiurnal tidal systems. During ebb stage, sand deposition takes place only in trough of sand wave and is preserved in form of intrasets within tosets of cross-bed set (C). During stillstand associated with turning of tide, clay accumulates on lee face and within trough of sand wave and is preserved as mudstone drapes (B, D, F).

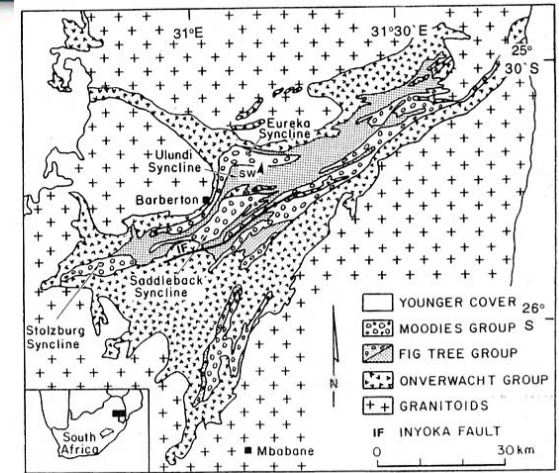


Figure 1. Simplified geological map of Barberton Greenstone Belt. Heavy arrowhead indicates location of sand-wave (sw) in Eureka syncline.

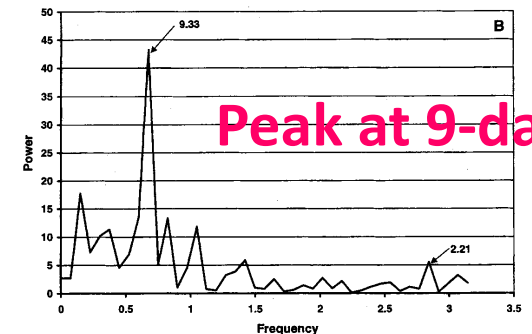
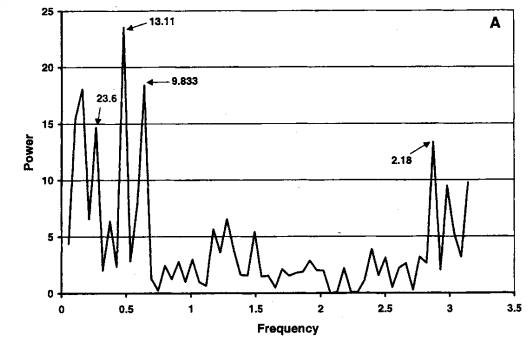


Figure 5. A: Traverse two—all data. Power spectral plot of foreset bundle thicknesses measured along traverse two (see Fig. 3A). B: Traverse two—subordinates removed. Power spectral plots of dominant flood-tide foreset bundle thicknesses along traverse two (see Fig. 3B).

Peak at 9-days

Komatiite: A first oceanic crust from a hot mantle!



Canada (Fig. 1.2) by Pyke et al. (1988). The lower chill contact, an overlying "hopper" olivine zone (B1) and a zone of bladed or plate spinifex which, in turn, is overlain by a zone of fine-grained flow top or "hopper" olivine zone. In all the komatiite flows in the Kambalda area, the flows are generally about 1 m thick and their thicknesses can be determined.



Fig. 1.2. Diagrammatic section of a typical komatiite flow.

Locality 2. Approximately 30m north of the Kambalda area, exposures of komatiitic pillow basalts are seen in a zone about 10 m wide. The pillow structures are formed as the lavas were erupted into water.

Locality 3. About 50 m north of locality 2, exposures of komatiitic pillow basalts (Fig. 1.3) are seen in a view on the flat pavements and in section on the pavements and show smooth curved upper surfaces.



Dyke and Komati River



Komatiite!! Please see my sample!



Komatiitc basalt: Pillow Lava



Question:
What does these
Komatiite mean?

<http://aginnoiators.org.au/sites/default/files/Spinifex>
(Triodia) grass in NT

**High temp. (1600°C) Crust
mantle**

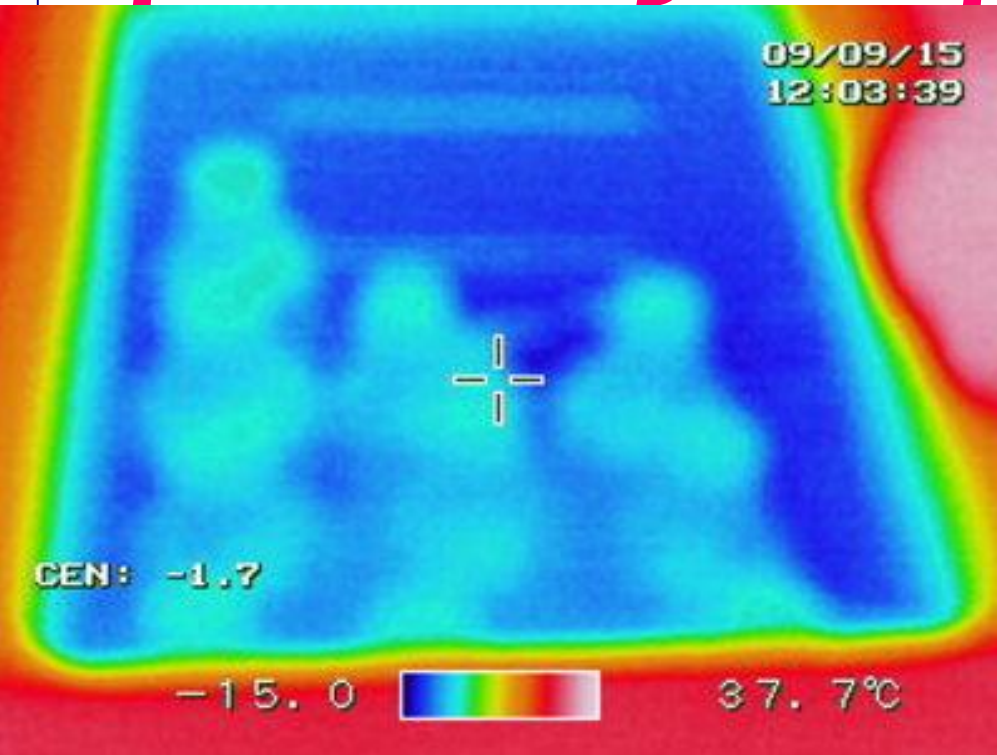
Super Cooling: Spinifex texture

First Oceanic Crust!



High temp. (1600°C) of early mantle

Super cooling -> Spinifex texture



**Special thanks to
Prof. Konishi
(Osaka Kyoiku University)**

After LHB a large impact still recorded on South Africa

LETTER

doi:10.1038/nature10982

Impact spherules as a record of an ancient heavy bombardment of Earth

B. C. Johnson¹ & H. J. Melosh^{1,2}

Nature 2012

Impact craters are the most obvious indication of asteroid impacts, but craters on Earth are quickly obscured or destroyed by surface weathering and tectonic processes¹. Earth's impact history is inferred therefore either from estimates of the present-day impactor flux as determined by observations of near-Earth asteroids, or from the Moon's incomplete impact chronology²⁻⁴. Asteroids hitting Earth typically vaporize a mass of target rock comparable to the projectile's mass. As this vapour expands in a large plume or fireball, it cools and condenses into molten droplets called spherules⁵. For asteroids larger than about ten kilometres in diameter, these spherules are deposited in a global layer. Spherule layers preserved in the geologic record accordingly provide information about an impact even when the source crater cannot be found¹. Here we report estimates of the sizes and impact velocities of the asteroids

that created global spherule layers. The impact chronology from these spherule layers indicates that the impactor flux was significantly higher 3.5 billion years ago than it is now. This conclusion is consistent with a gradual decline of the impactor flux after the Late Heavy Bombardment.

$$D_{\text{imp}} = 17(t_r/\xi)^{(1/3)} \quad (1)$$

where D_{imp} is the impactor diameter in kilometres and ξ is an efficiency factor that conservatively ranges from 0.5 to 2 for typical asteroidal impact velocities on Earth. Additionally, t_r is the layer's reduced thickness in centimetres, defined as $t_r = 2f_{\text{sp}}t$ where t is the measured layer thickness and f_{sp} is the volume fraction of spherules in the layer.

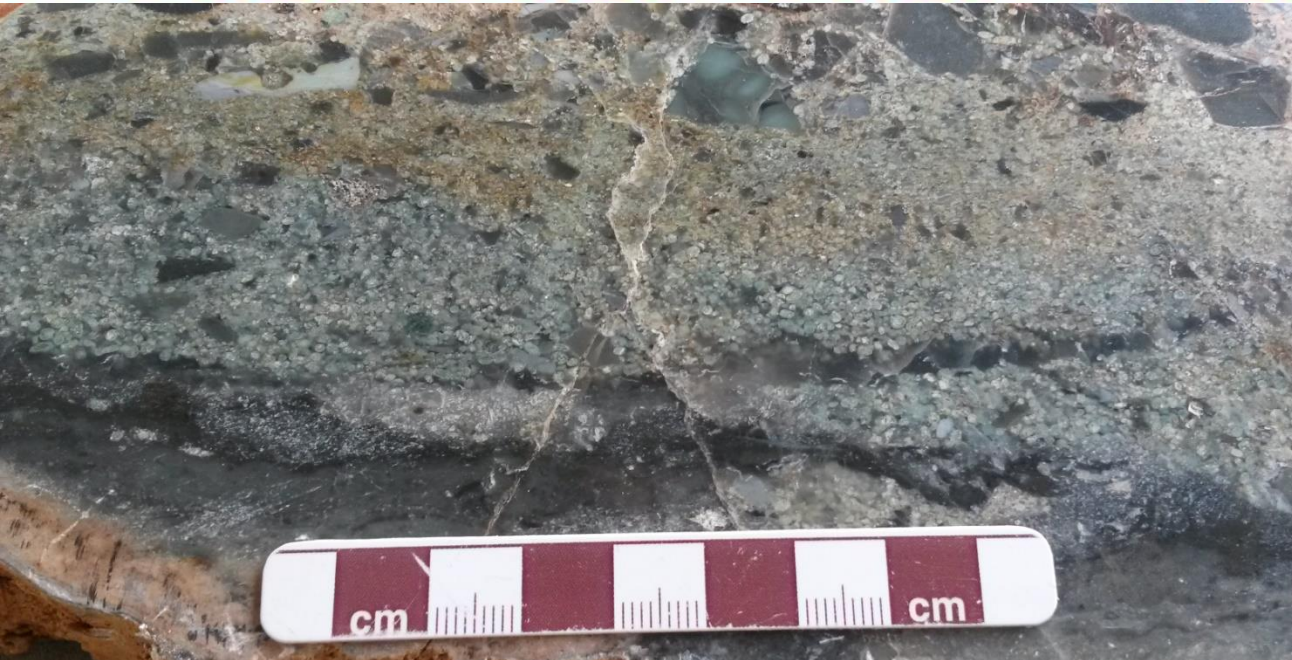
We test the accuracy of equation (1) by comparing the impactor size estimated using spherule layer thickness with the impactor size determined by other methods. The Cretaceous–Palaeogene boundary layer is found at numerous sites globally and has a thickness of around 3 mm and consists of about half spherules by volume⁶. Using the entire range of ξ and $D_{\text{imp}} = 9.0 \pm 1.1$ km, this is consistent with the size of the Chicxulub impactor (10 ± 4 km) as determined by iridium fluence and similar estimates from the size of the Chicxulub impact structure^{10,11}. Our impactor size estimates for two other spherule layers, S2 and S3 (Table 1), are also consistent with the estimate, 3–7 times larger than the Chicxulub impactor, which is also based on total iridium fluence¹².

The estimates of impactor sizes using equation (1) are only valid if the spherule layer is from a single impact event. The thickness of the global

Asteroid $\phi \sim 50$ km: crater size ~ 400 km
Hit a continent (now South Africa and Australia)



A Spherule sample at Barberton, South Africa (3.5–3.2Ga)



**Pictures by Tony Ferrar
Dr. Dion' s friend
He sent me these
pictures.**

**.Sample from Fig Tree G (3.4 Ga
Barberton)**



First life? Famous Schopf's paper 1993



University News

Oldest fossils controversy resolved

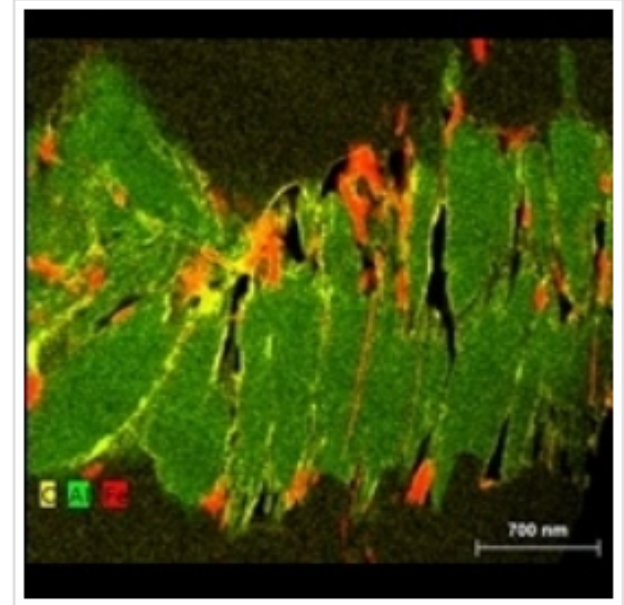
Monday, 20 April 2015

New analysis of world-famous 3.46 billion-year-old rocks by researchers from The University of Western Australia is set to finally resolve a long-running evolutionary controversy.

The new research, published this week in *Proceedings of the National Academy of Sciences USA*, shows that structures once thought to be Earth's oldest microfossils do not compare with younger fossil candidates but have, instead, the character of peculiarly shaped minerals.

In 1993, US scientist Bill Schopf described tiny (c. 0.5-20 micrometres wide), carbon-rich filaments within the 3.46 billion-year-old Apex chert from the Pilbara region of Western Australia, which he likened to certain forms of bacteria, including cyanobacteria. (Chert is fine-grained, silica-rich sedimentary rock.)

These 'Apex chert microfossils' soon became enshrined in textbooks, museums displays, popular science books and online reference guides as the earliest evidence for life on Earth. In 1996, these structures were even used to test and help refute the case against 'microfossils' in the Martian meteorite ALH 84001.



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A Bio-mat sample at Barberton, South Africa (3.2 Ga)

nature

Vol 463 | 18 February 2010 | doi:10.1038/nature08793

LETTERS

Organic-walled microfossils in 3.2-billion-year-old shallow-marine siliciclastic deposits

Emmanuelle J. Javaux¹, Craig P. Marshall² & Andrey Bekker³

Although the notion of an early origin and diversification of life on Earth during the Archaean eon has received increasing support in geochemical, sedimentological and palaeontological evidence, ambiguities and controversies persist regarding the biogenicity and syngeneity of the record older than Late Archaean^{1–3}. Non-biological processes are known to produce morphologies similar to some microfossils^{4,5}, and hydrothermal fluids have the potential to produce abiotic organic compounds with depleted carbon isotope values⁶, making it difficult to establish unambiguous traces of life. Here we report the discovery of a population of large (up to about 300 μm in diameter) carbonaceous spheroidal microstructures in Mesoarchaean shales and siltstones of the Moodies Group, South Africa, the Earth's oldest siliciclastic alluvial to tidal-estuarine deposits⁷. These microstructures are interpreted as organic-walled microfossils on the basis of petrographic and geochemical evidence for their endogenicity and syngeneity; their

the base of the Moodies Group, in interlayered laminated grey shales, siltstones and wavy-laminated clay-rich and organic-matter-rich layers, possibly representing microbial mat structures. Flaser bedding, small-scale cross-bedding, and mud-draped current ripples were observed in drill core samples, polished slabs and thin sections (Supplementary Fig. 2). These sedimentary structures indicate deposition in shallow-water environments above the wave base.

The Moodies Group is the uppermost of three stratigraphic units that comprise the Swaziland Supergroup in the BGB (Supplementary Fig. 1b). It consists of an up to 3.7-km-thick succession of alluvial to shallow-marine sandstones with subordinate conglomerates and mudstones, as well as iron formation and volcanic rocks¹⁵. Deposition of the Moodies Group began shortly after 3,226 ± 1 and 3,222 ± 10/–4 Myr ago (age of an ignimbrite and porphyritic intrusion, respectively, at the top of the underlying Fig Tree Group^{15,16}) but before 3,207 ± 2 Myr ago (age of a dacitic dyke cross-cutting the basal part of the Moodies



Javaux et. al., Nature 2010



Banded iron Formation (Sample)

Oxygen in the Ocean
Photo-synthesis by Cyanobacteria



Coffee Brake, at the Kruger national park, South Africa



Coffee Brake, at the Kruger national park, South Africa



One scene, at the Kruger national park, South Africa

- We enjoyed the game drive in the Kruger National Park
- Our driver opened the both side doors to view and made us to take picture easily ---.
- Unfortunately we came across a **patrol car**, we got stopped.
- **Two young black police men** walked to our car, One police man said to our driver Dr. Dion, “ **You commit a traffic violation**. The rule prohibit driving car with the door open. Opening the door is very dangerous. Because the wild animals running into the car! ”
- Then we all asked to the police “Please forgive our violation. Could you just let him off this one time?”
- Of course their answer was “**NO!**”. Our driver had ticketed.
- However, at that time **I was deeply moved this scenery!** -----.

Question:

Why was I deeply moved at that time?



Questions

What does the Komatiite mean?

At the Kruger N.P: Why was I deeply moved?

At the platinum mine: Why Is Japan the most important country as a customer?

Why are chromium layers so large and extended?

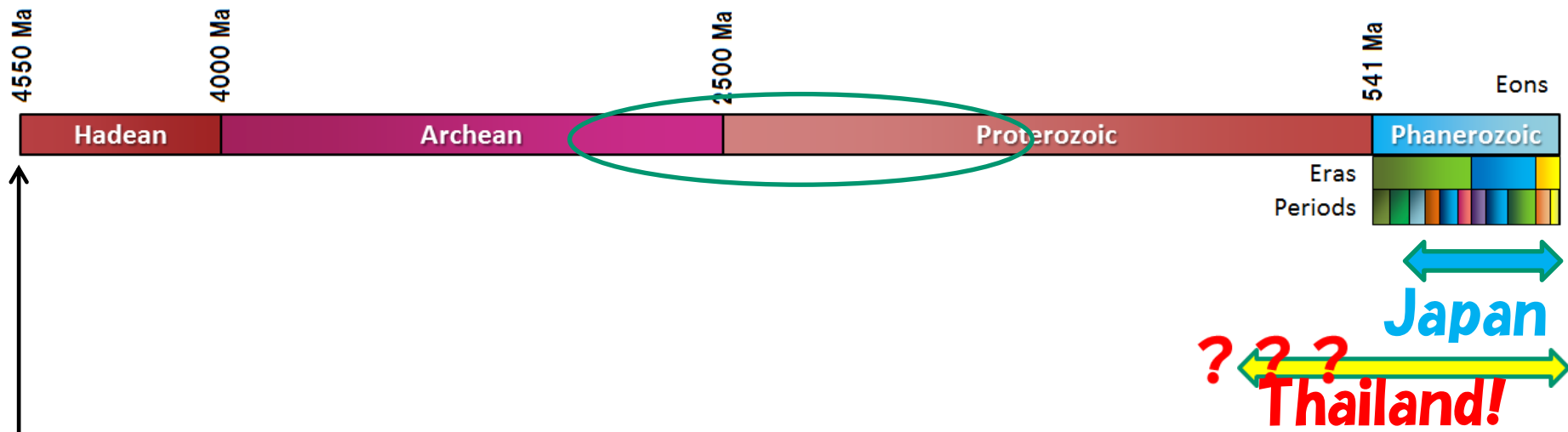
At the Gold mine: Why does the gold ore include pyrite(fool' s gold).

At the Impact Crater: At first how was the crater formed?

Why do I show this photo?



Early Earth (Part 3) Archean to Proterozoic



Gold deposit 3.0Ga

South Africa!

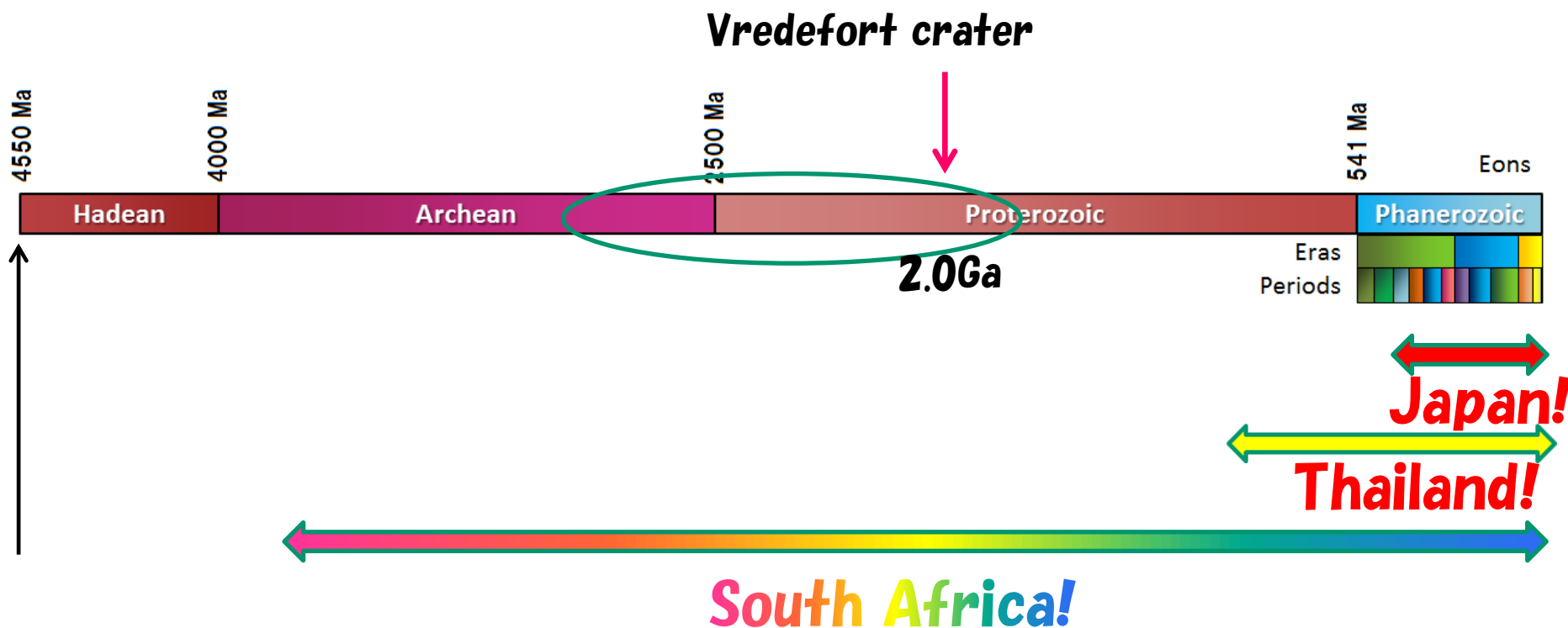
Chromium and Platinum: Bushveld Intrusion 4.5Ga

Diamond 4.4-4.0 Ga

Coal 0.35 Ga



Early Earth (Part 3) Archean to Proterozoic



Gold deposit 3.0Ga

Chromium and Platinum: Bushveld Igneous Complex (BIC) 2.0Ga

Vredefort impact crater 2.0Ga

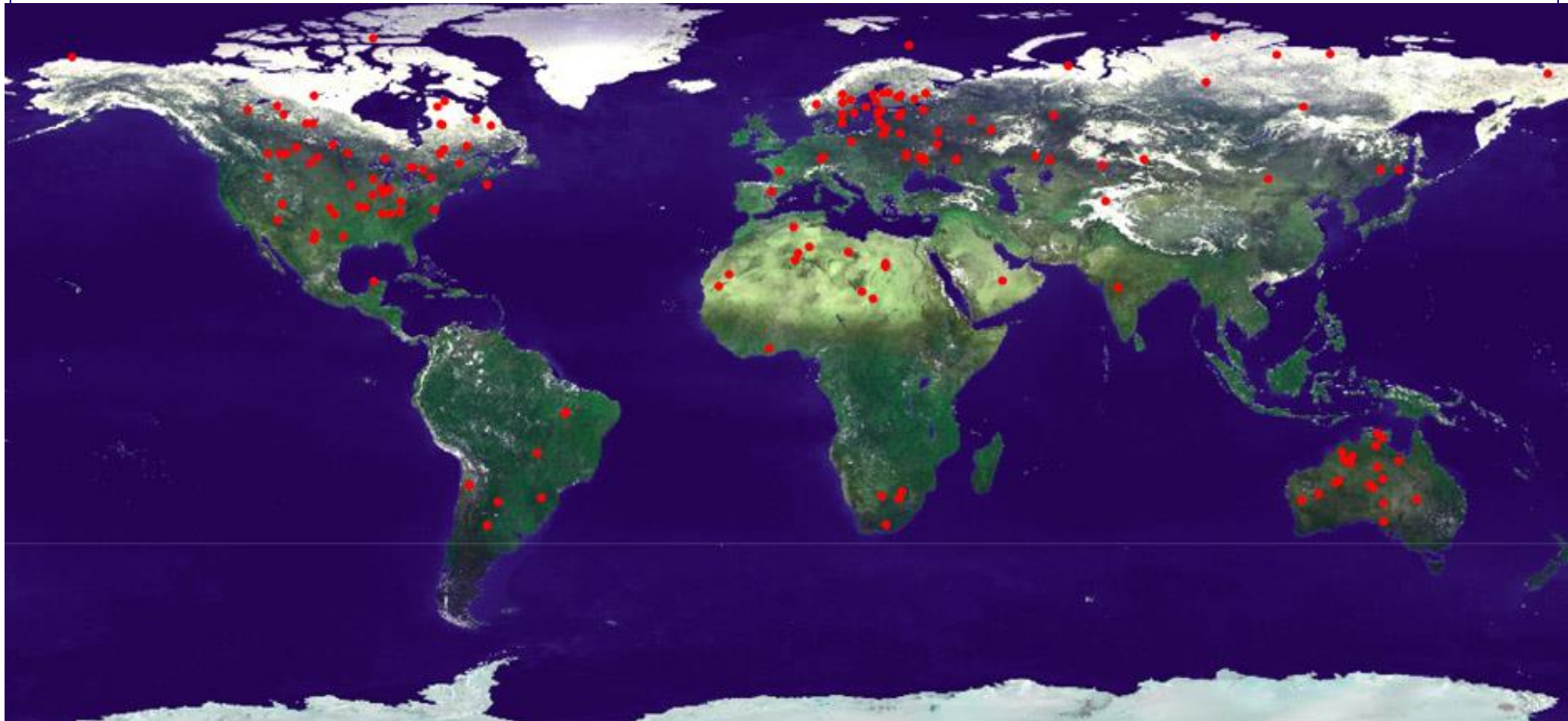
Diamond Kimberlite 1.2-0.1 Ga

(Coal 0.25 Ga)

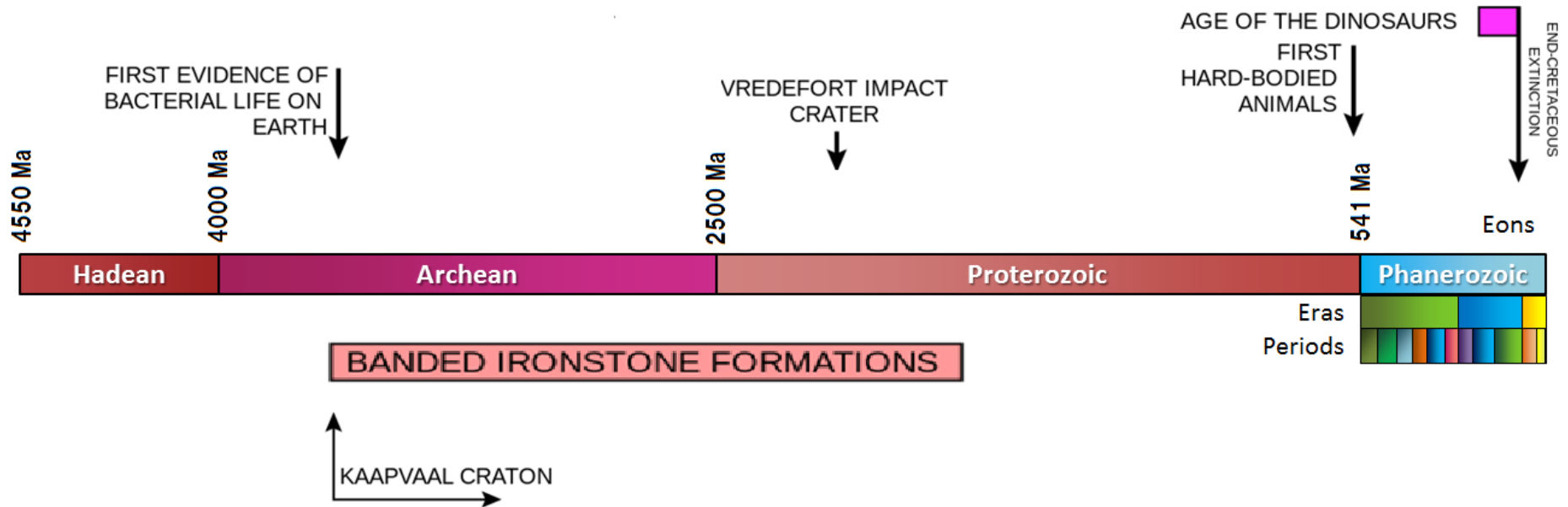


Impact Crater and Ore Deposits

http://www.lpi.usra.edu/science/Kring/epo_web/impact_cratering/World_Craters_web/worldcraters_maps.jpg

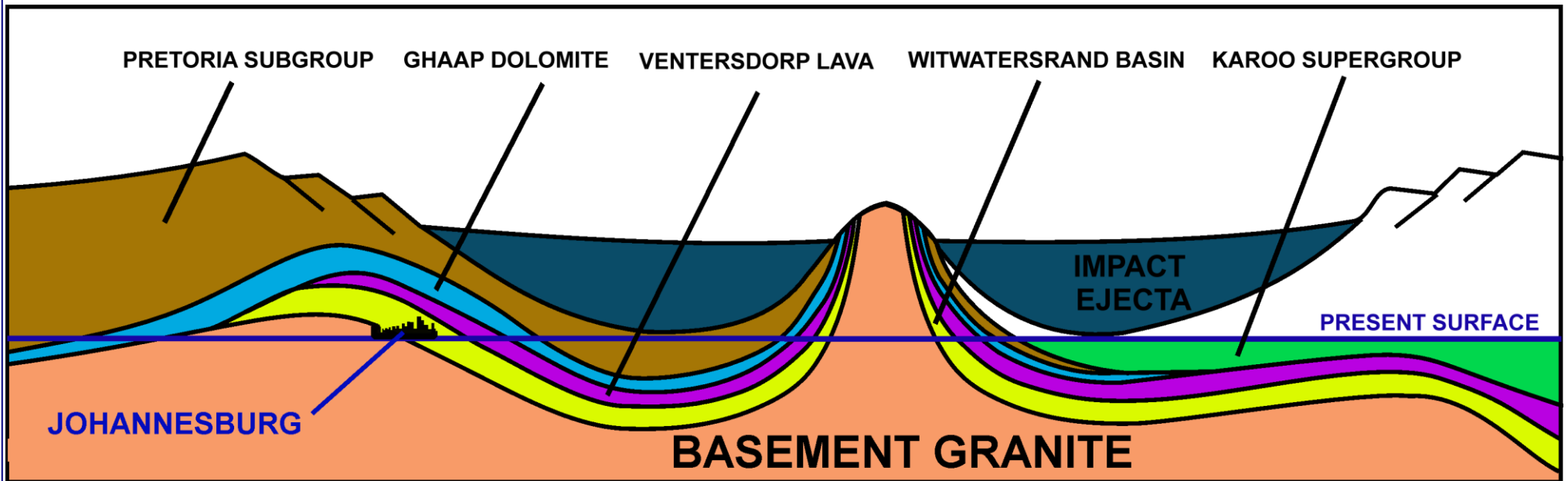


From wiki



A timeline of the earth's history indicating when the Vredefort crater was formed in relation to some of the other important South African geological events. W indicates when the Witwatersrand Supergroup was laid down, C the Cape Supergroup, and K the Karoo Supergroup. The graph also indicates the period during which banded ironstone formations were formed on earth, indicative of an oxygen-free atmosphere. The Earth's crust was wholly or partially molten during the Hadean Eon. One of the first microcontinents to form was the Kaapvaal Craton, which is exposed at the center of the Vredefort Dome, and again north of Johannesburg.





A schematic diagram of a NE (left) to SW (right) cross-section through the 2020 million year old Vredefort impact crater and how it distorted the contemporary geological structures. The present erosion level is shown. Johannesburg is located where the Witwatersrand Basin (the yellow layer) is exposed at the "present surface" line, just inside the crater rim, on the left. Not to scale.





300 km wide crater



Gold from Wikipedia

The gold in the Witwatersrand Basin area was deposited in Archean river deltas having been washed down from surrounding gold-rich greenstone belts to the north and west. Rhenium-osmium isotope studies indicate that the gold in those mineral deposits came from unusual 3000 million year old mantle-derived intrusions known as komatiite, present in the greenstone belts.

South Africa accounted for 15% of the world's gold production in 2002 and 12% in 2005, though the nation had produced as much as 30% of world output as recently as 1993. Despite declining production, South Africa's gold exports were valued at \$3.8 billion USD in 2005. The US Geological Survey estimated in that as of 2002, South Africa held about 50% of the world's gold resources, and 38% of reserves.



Tau Tona Gold Mine (Video)

It is one of the most efficient mines in South Africa and remains in continuous operation even during periods when the price of gold is low. Since its construction, two secondary shafts have been added bringing the mine to its current depth. The mine today has some 800 km (500 mi) of tunnels and employs around 5,600 miners. The mine is a dangerous place to work, with an average of five miners dying in accidents each year. The mine is so deep that temperatures in the mine can rise to life-threatening levels. Air conditioning equipment is used to cool the mine from 55 ° C (131 ° F) down to a more tolerable 28 ° C (82 ° F). The rock face temperature currently reaches 60 ° C (140 ° F).

Movie please!!



Fool's Gold!! (Why Pyrite?)

Not gold (Au), but pyrite (FeS_2)

3.0 Ga **reducing environment** of earth surface.

No Oxygen atmosphere, now our surface is covered by **iron-oxide**, hematite or magnetite.



Diamonds (Video)



Ever since the **Kimberley diamond** strike of **1868**, South Africa has been a world leader in diamond production. The primary South African sources of diamonds, including seven large diamond mines around the country, are controlled by the **De Beers** Consolidated Mines Company. In **2003**, De Beers operations accounted for **94%** of the nation's total diamond output of **11,900,000 carats (2.38 t)**. This figure includes both gem stones and industrial diamonds. Diamond production rose in **2005** to over **15,800,000 carats (3.16 t)**.



Platinum (Video) and palladium

South Africa produces more platinum and similar metals than any other nation. In 2005, 78% of the world's platinum was produced in South Africa, along with 39% of the world's palladium. Over 163,000 kilograms (5,200,000 ozt) of platinum was produced in 2010, generating export revenues of \$3.82 billion USD. Palladium is produced in two ways: recovery and mining production. Currently Russia and South Africa are the biggest palladium producers in the world.



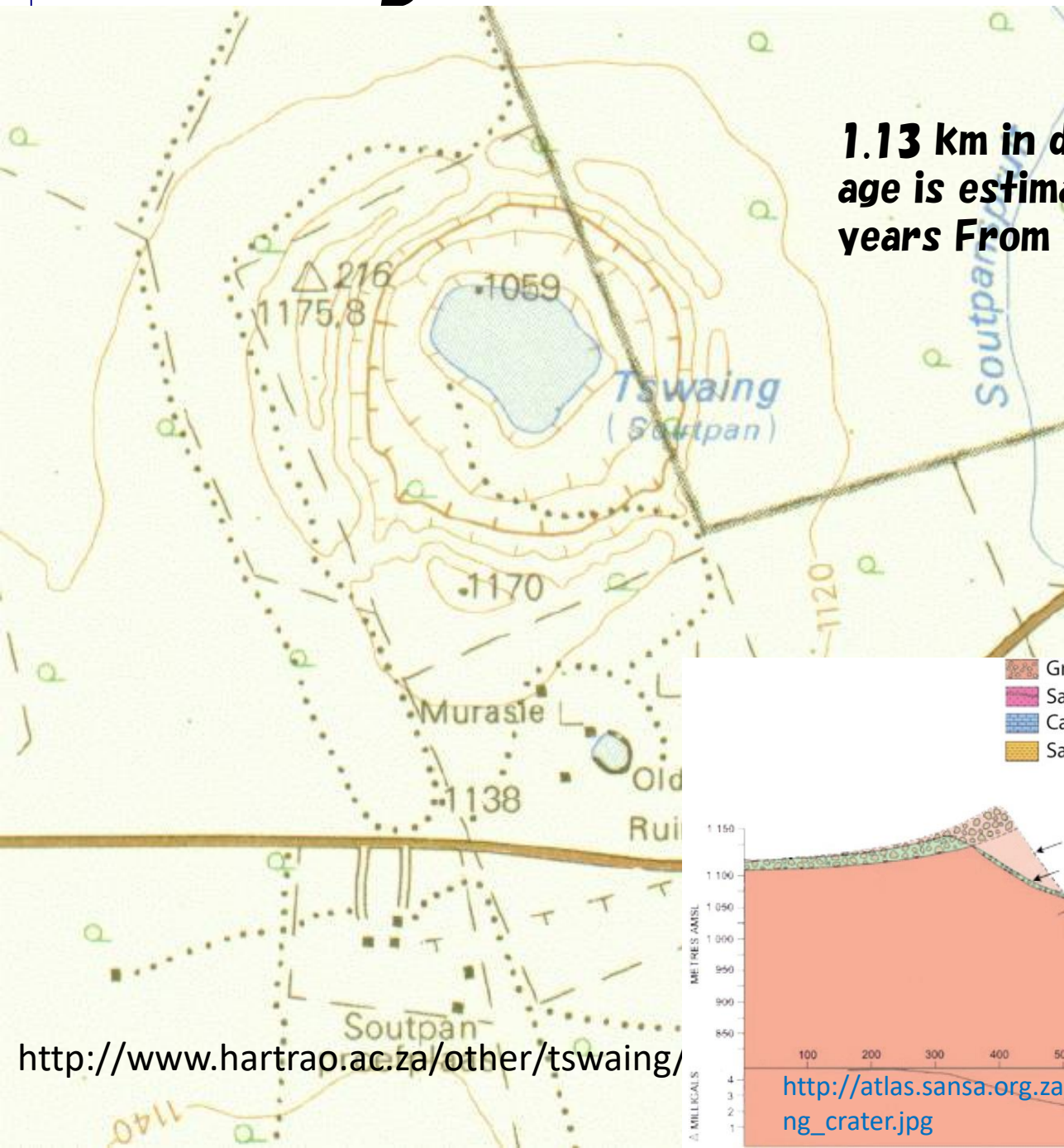
Chromium (Video)

Chromium is another leading product of South Africa's mining industry. The metal, used in stainless steel and for a variety of industrial applications, is mined at 10 sites around the country. South Africa's production of chromium accounted for 100% of the world's total production in 2005, and consisted of 7,490,000 metric tons (7,370,000 long tons; 8,260,000 short tons) of material.

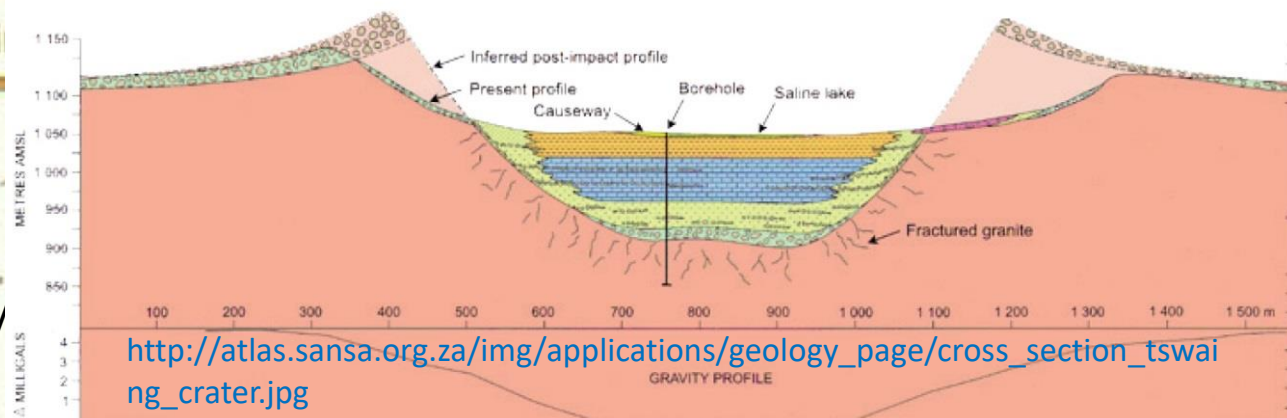


Tswaing Meteor Crater

1.13 km in diameter and 100 m deep and the age is estimated to be 220,000 ± 52,000 years From Wikipedia.

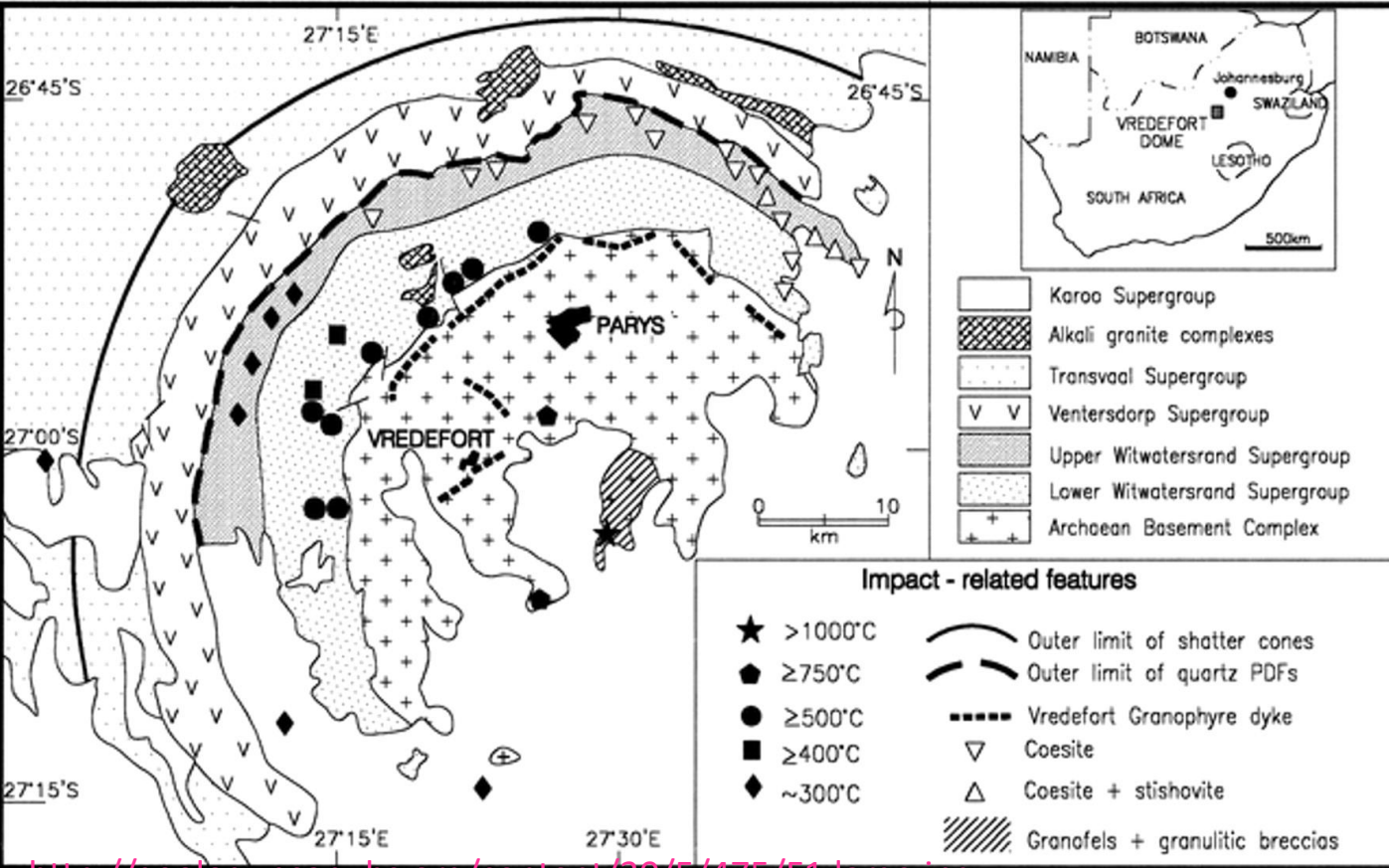


- Granite breccia
- Sand and boulders
- Carbonate rich sediments
- Saline muds
- Colluvium
- Colluvium fan
- Talus
- Fragmental granite breccia



<http://www.hartrao.ac.za/other/tswaing/>

http://atlas.sansa.org.za/img/applications/geology_page/cross_section_tswaing_crater.jpg



Vredefort crater
 more than 300 km across
 2.023 billion years
 (\pm 4 million years),

<http://geology.gsapubs.org/content/30/5/475/F1.large.jpg>





Photo by H.Nemoto (2010)

The Rainbow Country

-Peoples and cultures-



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14th July 2011 at the English Reading Class



PSI entrance and Dr. Hartmann

Life origin?

Lineweaver & Norman, 2008

When and How life appeared

Magma ocean and dry up

Sterilizing Impacts and life

“Panspermia Hypothesis”

A cool early Earth (John W. Valley)

April 2002: v. 30: no. 4: p. 351-

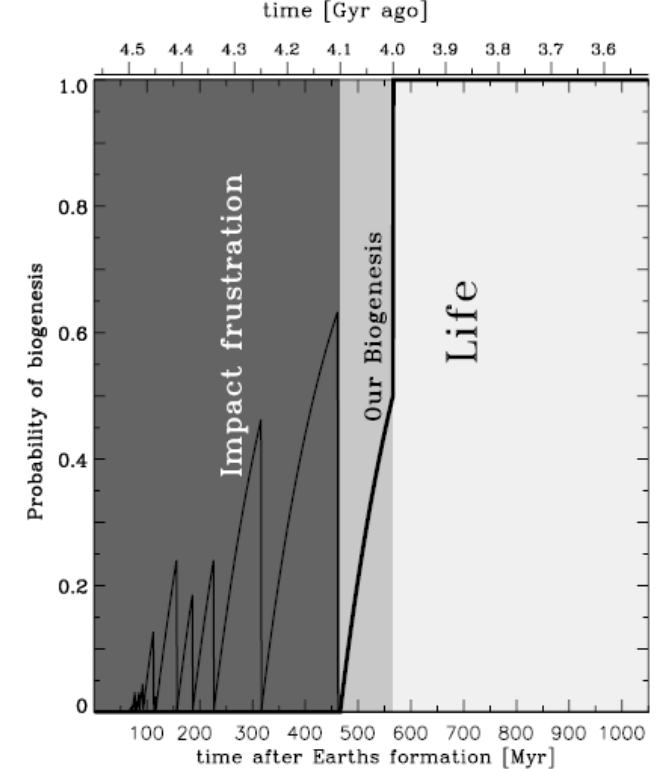
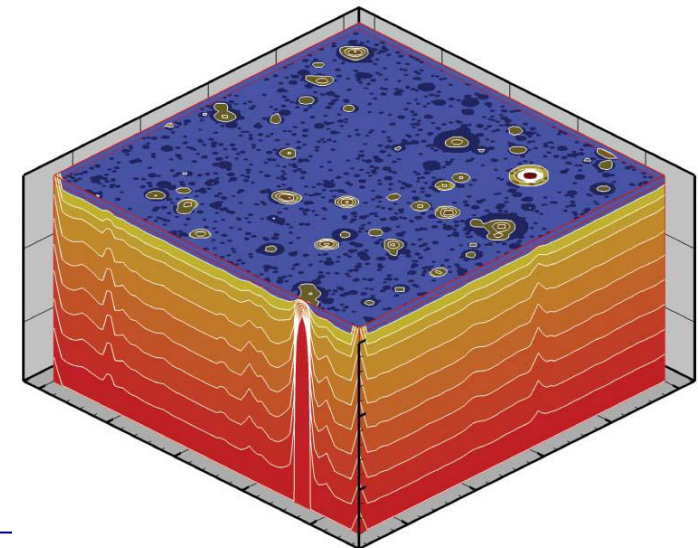
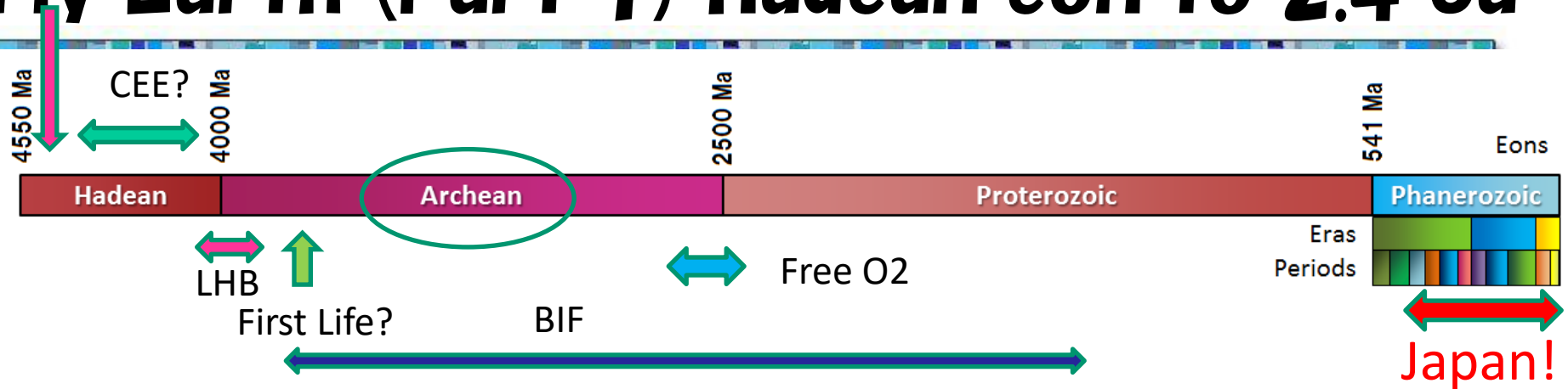


Fig. 2. After the sterilizing impact that formed the Moon about 90 ± 20 Myr after the formation of the solar system (Halliday 2008), a heavy but decreasing and stochastic bombardment lasted for a few hundred million years probably frustrated the origin of life on Earth. Eventually, the molecular evolution that led to life as we know it, was able to squeeze through the thermal bottlenecks produced by impacts (however see Abraomov & Mojzsis 2008a,b). Figure from Davies & Lineweaver 2005.



Abramov & Mojzsis, 2009

Early Earth (Part 1) Hadean eon to 2.4 Ga



Forming of Earth 4.6Ga

Giant Impact (the birth of Moon) 4.5Ga

Cool Early Earth 4.4-4.0 Ga

Late Heavy Bombardment (LHB) 3.9 Ga

